

Multi-messenger real-time analysis framework of the KM3NeTneutrino telescope

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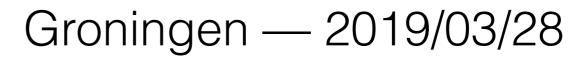
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Multi-messenger real-time analysis framework of the KM3NeT neutrino telescope

D. Dornic (CPPM/CNRS)

On behalf the KM3NeT Collaboration





Context

KM3NeT

Open questions:

- Origin of high-energy cosmic rays: which sources? What acceleration mechanisms? Which source evolutions? (mysteries of UHECR ?)
- Origin of IceCube HE astrophysical neutrinos
- Disentangle astrophysical models with multi-messenger observations
- Study of galactic (and extra galactic) propagation of CR with neutrinos as tracers
- Test the neutrino sector of the SM and BSM physics

So far, GW170817, IC170922, ANT150901, etc have demonstrated the capabilities of doing real-time multi-messenger follow-ups:

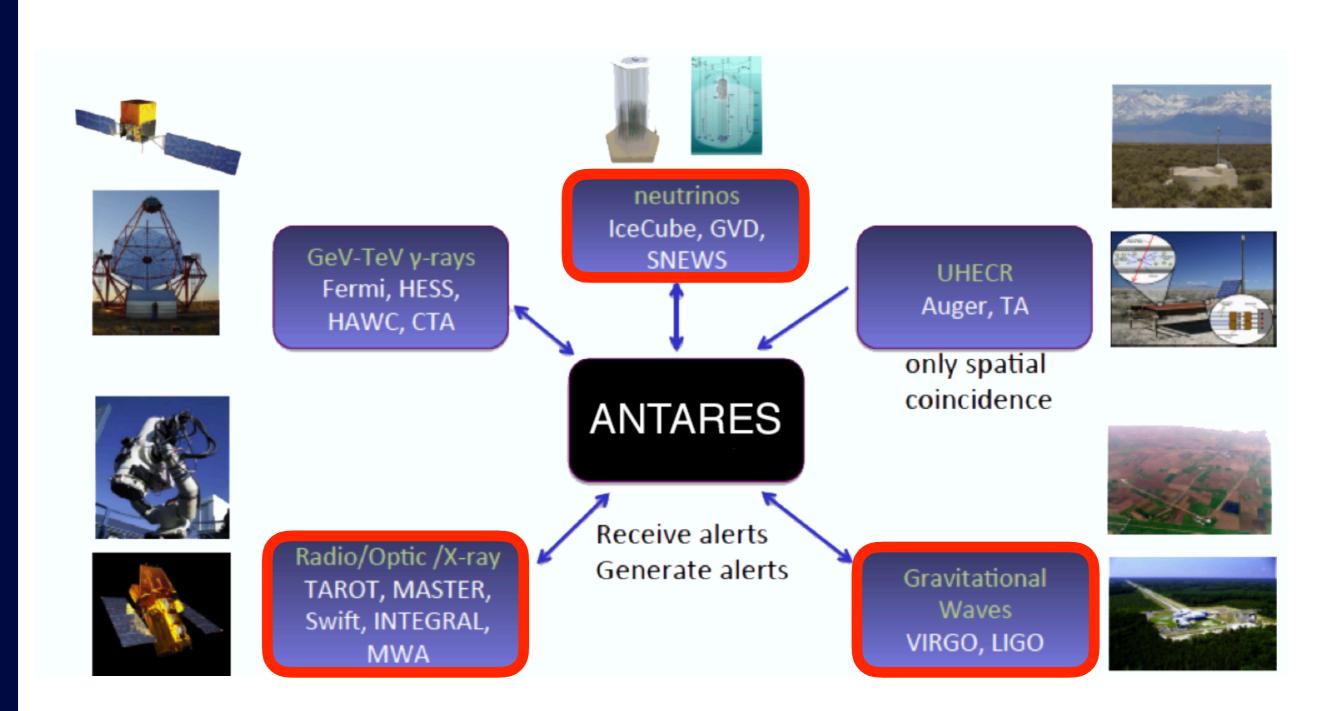
- Most of the HE sources are time-dependent with the flux quickly varying
- Provide accurate positions (required for redshift, host measurements)

- Maximize the scientific return of this event having a larger and more complete follow-up.

- Achieve simultaneous observations of transient phenomena by pointing instruments (so important for the modelisation)

- Determine the nature of a single event

Multi-messenger analysis



KM3NeT

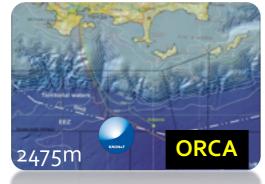
KM3NeT

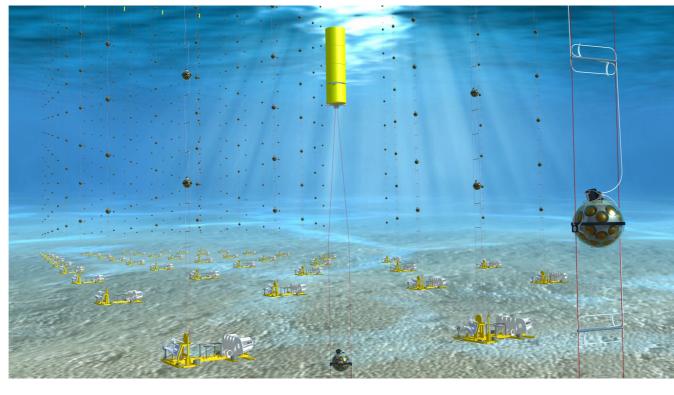


KM3NeT is the neutrino research infrastructure in the deep Mediterranean Sea



ORCA: off shore Toulon, France





Astroparticle Research with Cosmics In the Abyss

ARCA: off shore Capo Passero, Italy



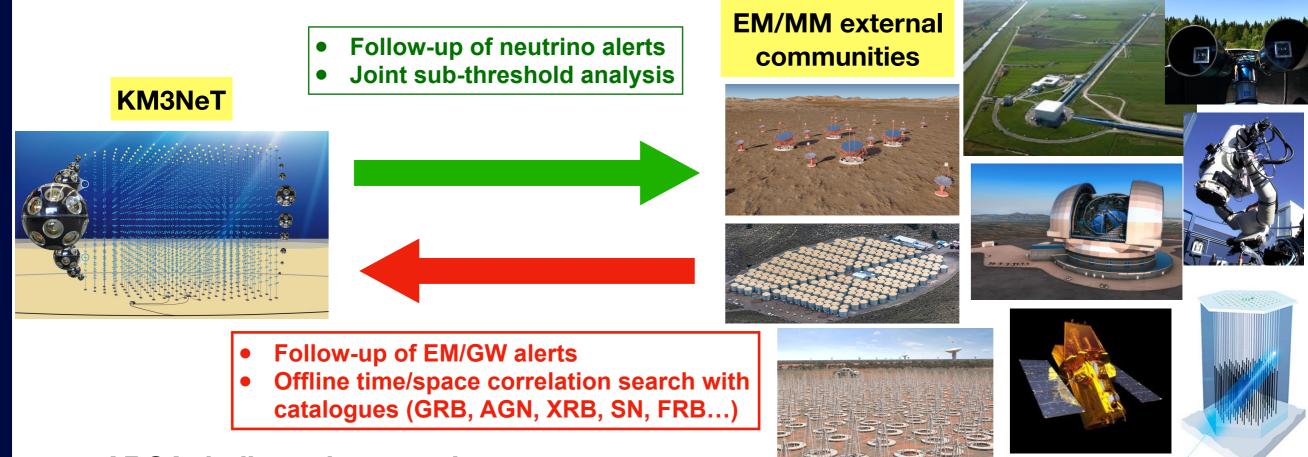
Main characteristics:

- Extended energy range: 3 GeV → 10 PeV (+ 10-40 MeV)
- Full sky coverage with the best sensitivity for the galactic sources
- High duty cycle (> 90-95%)
- All-flavour neutrino detection
- Good angular resolutions

 \implies Construction on-going: 1 DU working in ARCA & ORCA + 5 DUs ready for deployment in ORCA (+300 DOMs builded)

 \implies Mid 2020, better sensitivities than ANTARES in the whole energy range.

KM3NeT multi-messenger analyses

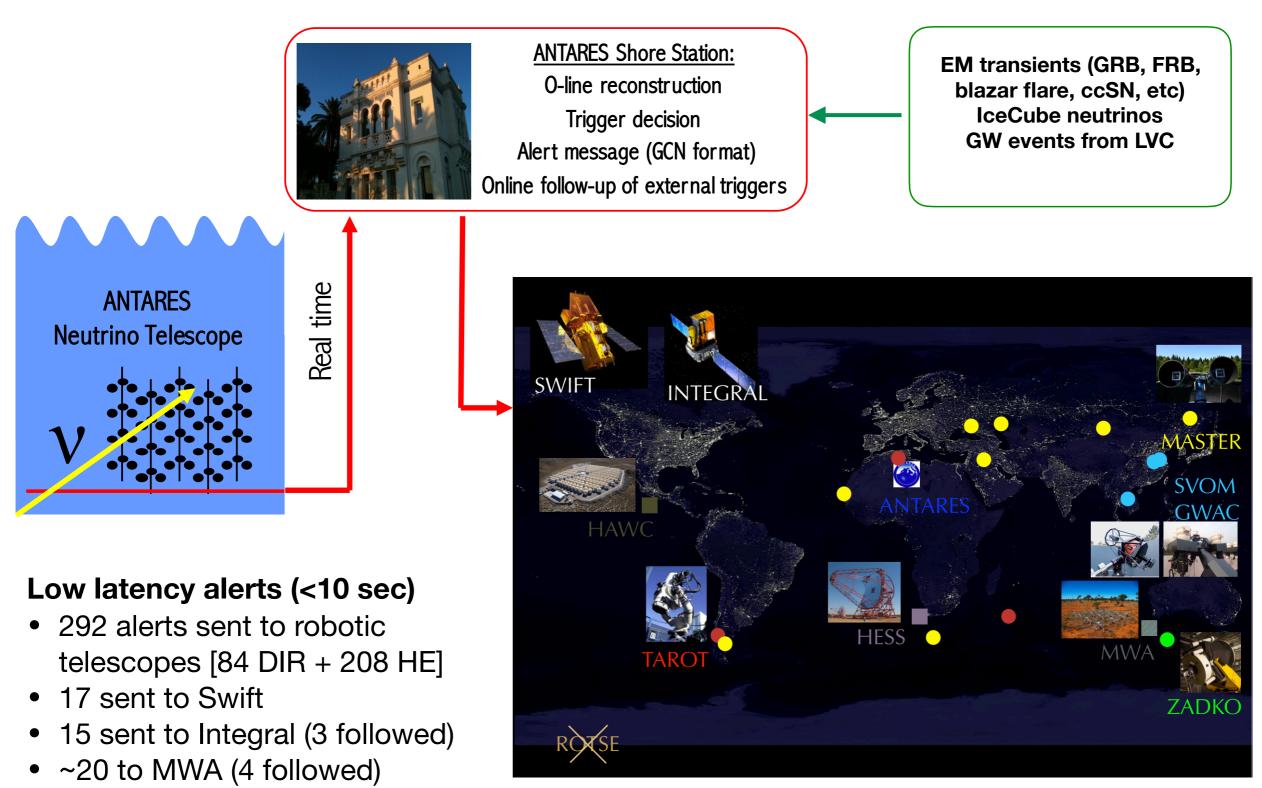


- ARCA dedicated to neutrino astronomy:
 - \implies Tracks (100 TeV 10 PeV) with the excellent angular resolution (<0.2°)
- \implies Cascades (100 TeV 10 PeV) thanks to the good angular resolution (1-2°) taking the advantage of the low atmospheric background contribution

- ORCA can do also astronomy:

- \implies Tracks & cascades at low energy (few GeV 10 TeV), looking for time/space clusters
- \implies Example sources: winds of binaries, chocked GRBs, hidden jets in core-collapse SN
- ORCA & ARCA: detection of MeV neutrinos from core-collapse SN

ANTARES online framework

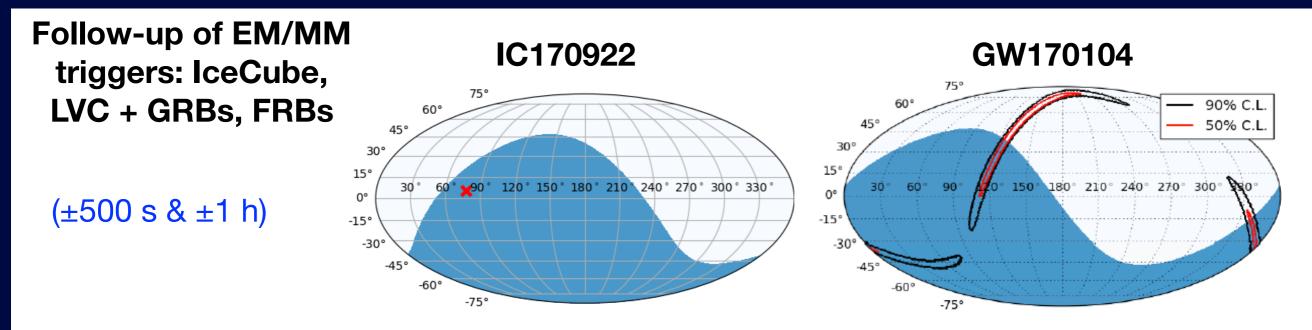


• 2 to HESS

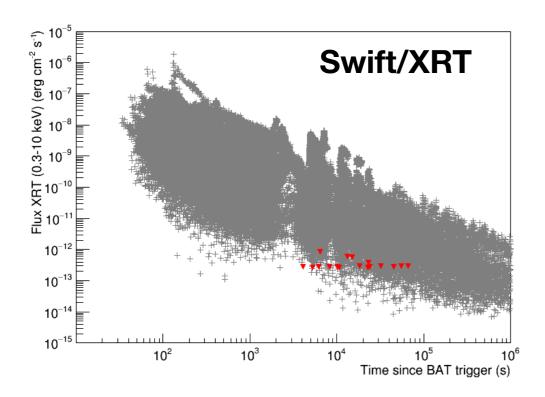
KM3Ne

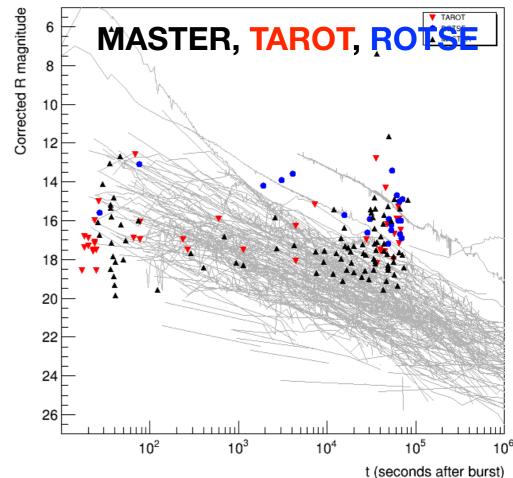
Examples of online ANTARES analyses





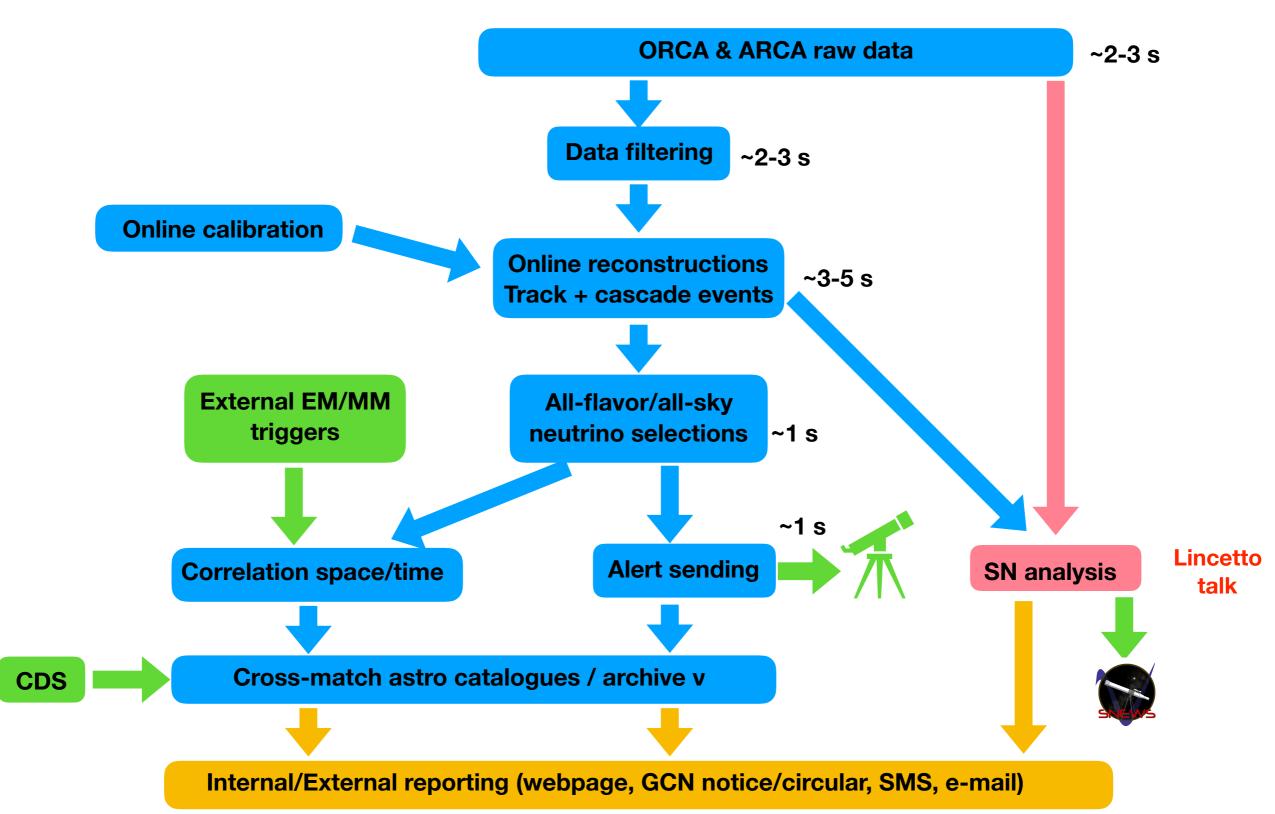
Follow-up of ANTARES alerts





7

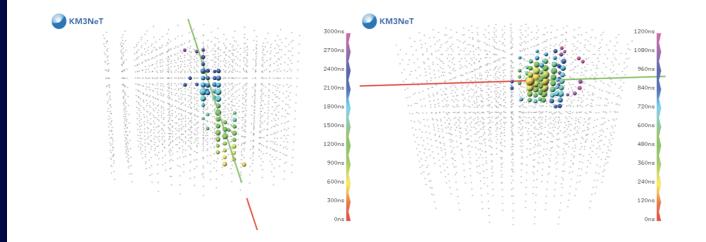
KM3NeT real-time framework



KM3NeT

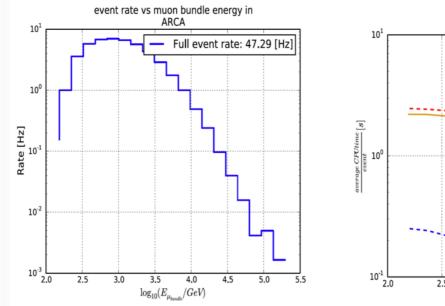
Online event reconstructions

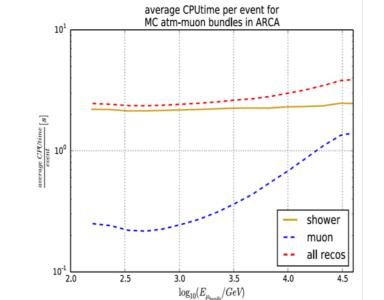
* All-flavor (track+cascade) event reconstructions: same framework and the same reconstruction tools as in offline



<u>Tracks</u>: ARCA: < 0.2° (>10 TeV) ORCA: 1 - 2° (100 GeV - 1 TeV) <u>Cascades</u>: ARCA: 1.5 - 2° (>10 TeV) ORCA: ~4 - 5° (100 GeV - 1 TeV)

* Time to reconstruct all events: Trigger rate: ~100 Hz \implies Neutrino rate: 1-2 mHz





SHOWER;	
$90 \text{TeV} < \text{E}_{\text{v}} < 110 \text{TeV}$	2.30 sec/event
900TeV < E _v < 1100TeV	2.80 sec/event
TRACKS	6
$\frac{\text{TRACKS}}{90\text{TeV}} < E_v < 110\text{TeV}$	S 0.85 sec/event

\implies Need 2 farms of 200 CPUs

KM3Ne¹

Sending alert system

Alert sending policy:

- ➡ Typical alert rate: few per month
- Standard alerts will be distributed through private channel to observing teams upon MoU agreements like ANTARES.
- After a commissioning phase, notable events will trigger alerts that will be distributed publicly to the astro community [Open Public Alert program]

Alert distribution:

- Distribution via the GCN network
- Message: VO event (XML file)
- Only 2 brokers for public and private alerts for both KM3NeT detectors

Reporting:

- SMS/e-mail to alert KM3NeT shifters
- Automatic GCN notices in case of very interesting neutrino signals
- KM3NeT subgroup shifters (check detector stability, update reconstructions, etc)
- ➡ GCN circular sent for refined information or identified counterpart (+ retraction).
- Results displayed in public/internal webpages

KM3NeT neutrino alerts

ANTARES alert distributions:	
* GCN socket: TAROT, ZADKO, MASTER, INTEGRAL * VO Event: MWA, HESS, SVOM, AMON	
* Mail: Swift	
For ANTARES, neutrino information is private. Need MoU	
with external partners.	
· · · ·	
Alert Message: Only one real-time message	
* ID	
* Time,	
* RA, DEC, error 50%	
* Energy proxy	
* Reconstruction quality	
* probability neutrino	For KM3NeT:
* Multiplicity, type of trigger	* ID
	* Detector (C
	* Time
selected / KM3NeT / /	* RA, DEC,
neutrinos // / Broker / //	* Energy est
	* Reconstruct
ORCA Private	* Probability
	* Type of ne
	* Multiplicity
neutrinos Broker	
	* Type of trig
	+ develop on
(SN)→	ARCA with d

efine a standard VO event:

- RCA/ARCA)
- ror 50, 90%
- nate
- on quality
- eutrino
- rino
- er
- alert broker for ORCA & erent types of alerts

KM3NeT

Summary

KM3NeT

- Despite its small size, ANTARES has performed plenty of multi-messenger analyses with more than 10 years of data, some really competitive with IceCube. Existing experiences for setting KM3NeT multi-messenger program.
- By observing astrophysical neutrinos with an unprecedented angular resolution, an extended energy range and a full sky coverage, KM3NeT will play a key role.
- The construction of ORCA and ARCA is on-going. First data looks good and first data analysed to validate the detector performance.

 \implies Setting the data acquisition using standard tools (IVOA, ASTERICS, CDS) and prepare the multi-messenger analyses.