

# Search for charged massive long-lived particles

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#### A search for charged massive long-lived particles

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We report on a search for charged massive long-lived particles (CMLLPs), based on 5.2 fb<sup>-1</sup> of integrated luminosity collected with the D0 detector at the Fermilab Tevatron  $p\bar{p}$  collider. We search for events in which one or more particles are reconstructed as muons but have speed and ionization energy loss (dE/dx) inconsistent with muons produced in beam collisions. CMLLPs are predicted in several theories of physics beyond the standard model. We exclude pair-produced long-lived gaugino-like charginos below 267 GeV and higgsino-like charginos below 217 GeV at 95% C.L., as well as long-lived scalar top quarks with mass below 285 GeV.

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We report on a search for massive particles, which are electrically charged and have a lifetime long enough to escape the D0 detector before decaying. Charged massive long-lived particles are not present in the standard model (SM) nor are their distinguishing characteristics (slow speed, high dE/dx) relevant for most high energy physics studies. Although the distinctive signature in itself provides sufficient motivation for a search, some recent extensions to the SM suggest that CMLLPs exist and are not yet excluded by cosmological limits [1, 2]. Indeed, the standard model of big bang nucleosynthesis (BBN) has difficulties in explaining the observed lithium production. The existence of a CMLLP that decays during or after the time of BBN could resolve this disagreement [3].

We derive cross section limits for CMLLPs and compare them to theories of physics beyond the SM. In most supersymmetric (SUSY) models the lightest SUSY particle is assumed to be stable. Some SUSY models predict that the nextto-lightest supersymmetric particle (NLSP) can be a CMLLP. In this Letter we explore models which include a chargino as a NLSP whose lifetime can be long if its mass differs from the mass of the lightest neutralino by less than about 150 MeV[4, 5]. This can occur in models with anomalymediated supersymmetry breaking (AMSB) or in models that do not have gaugino mass unification. There are two general cases, where the chargino is mostly a higgsino and where the chargino is mostly a gaugino, which we treat separately in this Letter.

There are some SUSY models that predict a scalar top quark (stop) NLSP and a gravitino LSP. These stop quarks hadronize into both charged and neutral mesons and baryons which live long enough to be CMLLP candidates [6]. Further, hidden valley models predict GMSB-like scenarios where the stop quark acts like the LSP and does not decay but hadronizes into charged and neutral hadrons that escape the detector [7, 8]. In general, any SUSY scenario where the stop quark is the lightest colored particle (which will happen in models without mass unification and heavy gluinos) can have a stop CMLLP. Any colored CMLLP will undergo hadronization and charge exchange during nuclear interactions, which we discuss below.

This search utilizes data collected between 2006 and 2010 with the D0 detector [9] at Fermilab's 1.96 TeV  $p\bar{p}$  Tevatron Collider, and is based on 5.2 fb<sup>-1</sup> of integrated luminosity. We reported earlier [10] on a similar 1.1 fb<sup>-1</sup> study, searching

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for events with a pair of CMLLPs, each with low speed. In addition to using the larger data sample, the present search looks for one or more CMLLP, rather than only for a pair, and characterizes CMLLPs with high dE/dx in addition to slow speed. Other searches for long-lived particles include those from the CDF Collaboration [11, 12], the CERN  $e^+e^-$  Collider LEP [13], and the CERN pp Collider LHC [14, 15].

The D0 detector [9] includes an inner tracker with two components: an innermost silicon microstrip tracker (SMT) and a scintillating fiber detector. We find a particle's dE/dxfrom the energy losses associated with its track in the SMT. The tracker is embedded within a 1.9 T superconducting solenoidal magnet. Outside the solenoid is a uranium/liquidargon calorimeter surrounded by a muon spectrometer, consisting of drift tube planes on either side of a 1.8 T iron toroid. There are three layers of the muon detector: the A-layer, located between the calorimeter and the toroid, and the B- and C- layers, outside the toroid. Each layer includes scintillation counters which serve to veto cosmic rays. Thus the muon system provides multiple time measurements from which a particle's speed may be calculated.

Because we distinguish CMLLPs solely by their speed  $\beta$ and dE/dx, we must measure these values for each muon candidate as accurately as possible. Muons from  $Z \rightarrow \mu\mu$  events studied throughout the data sample allow calibration of the time measurement to better than 1 ns, with resolutions between 2-4 ns, and to maintain the mean dE/dx constant to within 2% over the data-taking period. From a specific muon scintillation counter we calculate a particle's speed from the time recorded and the counter's distance from the production point, and we compute an overall speed from the weighted average of these individual speeds, using measured resolutions. The ionization loss data from the typically 8-10 individual energy deposits in the SMT are combined using an algorithm that omits the largest deposit to reduce the effect of the Landau tail and corrects for track crossing angle. We normalize the dE/dx measurements by requiring the dE/dx distribution of muons from  $Z \rightarrow \mu \mu$  events to have a maximum at 1. Figure 1 shows the distributions in  $\beta$  and dE/dx for data and background, which pass the selection criteria described below.

The selection of a candidate CMLLP occurs in several steps. Because of the high  $p\bar{p}$  collision rate, we employ a three-level trigger system to reduce the event rate to the 200 Hz that can be recorded. The trigger system bases its decisions on characteristics of the event, which for the CMLLP candidates is the presence of a muon with a high momentum transverse to the beam direction  $(p_T)$ . In order to reduce triggers on cosmic rays, there is a time window at the initial trigger level. This trigger gate reduces the trigger efficiency by 10% for CMLLPs with a mass of 300 GeV (as they will be slow and some will be out-of-time) and contributes significantly to our overall acceptance. We avoid a tighter timing gate usually imposed at the second level of the muon trigger by accepting an alternative requirement that the muon have a matching track in the SMT.

In the standard D0 event reconstruction CMLLPs would ap-



Figure 1: Distributions of (a) speed  $\beta$  and (b) dE/dx for data, background, and signal (gaugino-like charginos with a mass of 100 and 300 GeV) that pass the selection criteria. The histograms have been normalized to have the same numbers of events. We have adjusted the scale of the dE/dx measurements so that the dE/dx of muons from  $Z \rightarrow \mu\mu$  events peak at 1. All entries exceeding the range of the histogram are added to the last bin.

pear as muons, which has been verified in detail using MC simulations. Thus we select events with at least one well identified high  $p_T$  muon. For a useful  $\beta$  measurement, the event must have scintillator hits in the A-layer and either the B- or C-layer; and for valid dE/dx data, we require at least three hits in the SMT. For an optimal tracking and momentum measurement we require the muon to be central, i.e., with a pseudorapidity [16]  $|\eta| < 1.6$ . To reject muons from meson decays we impose the isolation requirement that the sum of the  $p_T$  be less than 2.5 GeV for all other tracks in a cone of radius  $R = \sqrt{(\Delta \phi)^2 + (\Delta \eta)^2} < 0.5$ . We also require that the total transverse calorimeter energy in an annulus of radius 0.1 < R < 0.4 about the muon direction be less than 2.5 GeV.

A requirement that the z-coordinate of the muon track at the location of closest approach to the beam axis be < 40 cm ensures that the particle passes through the SMT.

We impose further criteria to eliminate cosmic rays. To select muons traveling outwards from the apparent interaction point, we require that its C-layer time be significantly greater than its A-layer time. We require also that the muon's distance-of-closest-approach to the beam line be less than 0.02 cm. These criteria are also applied to a second muon, if any, in the event. In addition, for events with two muons we require that the absolute value of the difference between each muon's A-layer times be less than 10 ns. To reject cosmics that appear as two back-to-back muons, we require for their pseudo-acolinearity  $\Delta \alpha = |\Delta \theta + \Delta \phi - 2\pi| > 0.05$ .

Events with a muon from a W boson decay, with mismeasurements providing inaccurate values of the muon's  $\beta$  and dE/dx, constitute a potentially large background. To study selection criteria for CMLLPs, we calculate the transverse mass  $M_T$  [17] and select data with  $M_T < 200 \mbox{ GeV}$  to model the data in the absence of signal [18]. We choose selection criteria that minimize the number of events surviving from this background sample compared to events from simulations of the CMLLP signal. We require that events contain at least one muon with  $p_T > 60$  GeV. From a separate sample of muons from  $Z \rightarrow \mu \mu$  decays we observe that the association of a spurious scintillator hit can result in an anomalously low  $\beta$ value. We use an algorithm that discards such hits through minimizing the  $\chi^2/d.o.f.$  for the  $\beta$  calculated from the different scintillator layers. By comparing the effect on the background sample with the effect on simulated signal, we choose to eliminate events unless the minimized speed  $\chi^2/d.o.f. <$ 2. Finally, we compare the muon's track direction measured by the muon system with that measured in the central tracker, and eliminate events with clearly mismatched tracks.

To simulate signal events, we generate CMLLP candidates using PYTHIA [19]. The long-lived stop quarks are hadronized using [20]. Because the signature of the CMLLP cascade decays is model dependent and difficult to simulate accurately, we generate direct pair-production of the CMLLPs, without including cascade decays. We use the full D0 detector GEANT [21] simulation to determine the detector response for these samples (which include overlaid databased  $p\bar{p}$  interactions). The results are applicable to models with pair-produced CMLLPs with similar kinematics.

The stop quarks are distinct since they appear in charged or uncharged stop hadrons, which may flip their charge as they pass through the detector. In the simulation approximately 60% of stop hadrons are charged following initial hadronization [22], i.e., 84% of the events will have at least one charged stop hadron. Further, stop hadrons may flip their charge through nuclear interactions as they pass through material. We assume that stop hadrons have a probability of 2/3 of being charged after multiple nuclear interactions and antistop hadrons, a probability of 1/2 of being charged, consistent with the numbers of possible stop and anti-stop hadronic final states [23–25]. For this analysis we require the stop hadron to be charged before and after passing through the calorimeter, i.e., to be detected both in the tracker and in the A-layer, and to be charged after the toroid, i.e., to be detected in the B- or C-layers. The probability for at least one of the stop hadrons being detected is then 38%, or 84% if charge flipping does not occur. We include these numbers as normalization factors in the confidence level analysis discussed below.

Our final selection criteria is that the candidate's speed  $\beta < 1$ . Thus, we describe the background by the  $\beta < 1$  data events with  $M_T < 200$  GeV, and search for CMLLP candidates in  $\beta < 1$  data with  $M_T > 200$  GeV. We normalize the background and data samples in the  $\beta > 1$  region, where the contribution of signal is negligible. Because the uncertainties in the speed measurements depend on the particle's  $\eta$ , due to detector geometry, and the distributions in  $\eta$  of the muons in the  $M_T < 200$  GeV sample differs from those in the  $M_T > 200$  GeV sample, we use the signal-free region to derive correction factors for the background sample that match its  $\eta$  distribution to that of the data.

We utilize a boosted decision tree (BDT) [26] to discriminate signal from background. The most discriminating variables are the CMLLP candidate's  $\beta$  and dE/dx, but we also include several related variables: the speed significance, defined as  $(1 - \beta)/\sigma_{\beta}$ , the corresponding number of scintillator hits, the energy loss significance defined as  $(dE/dx - 1)/\sigma_{dE/dx}$ , and the number of SMT hits. For each mass point in all three signal models we train the BDT with the signal MC and the background, and then apply it to the data samples. Figure 1 shows the distributions in  $\beta$  and in dE/dx for the data and background samples, as well as for two representative signals.

Systematic uncertainties are studied by applying variations to the background and signal samples and determining the deviations in the BDT output distributions. Two of the systematic uncertainties affect the shape of the BDT distribution and their effect is taken into account explicitly in the limit calculation: the uncertainty due to the width of the Level 1 trigger gate and the uncertainty in the corrections to the MC time resolution; both are applied only to signal. By examining the signal-like region of the BDT distributions, we find that the maximum (average) uncertainty is 10% (4%) for the trigger gate width, and 38% (7%) for the time resolution correction. All other systematic uncertainties affect only the normalization of the BDT output. The systematic uncertainties on the background are due to the dE/dx modeling (< 0.1%) and the background normalization, from the specific values used for the  $\beta$  (7.2%) and  $M_T$  requirements (2.2%). The systematic uncertainties on the signal include muon identification (2%) and the integrated luminosity (6.1%) [27]. The systematic uncertainties associated with the corrections to the muon  $p_T$  resolution and to the dE/dx resolution, as well as the choice of PDF and factorization scale, are all below 1%.

We obtain the 95% C.L. cross section limits from the BDT output distributions, constraining systematic uncertainties to data in background dominated regions [28]. These limits are shown in Figure 2, together with the NLO theoretical signal



Figure 2: 95% C.L. cross-section limits as a function of mass for (a) gaugino-like charginos, (b) higgsino-like charginos, and (c) stop quarks. The stop quark limits are displayed for the assumed charge flipping (charge survival probability = 38%) and for no charge flipping (charge survival probability = 84%).

cross sections, computed with PROSPINO [29]. The limits are of order 0.01 pb for directly pair-produced CMLLPs with masses between 200 and 300 GeV. Using the theoretical cross sections, we are able to exclude gaugino-like charginos below 267 GeV and higgsino-like charginos below 217 GeV. For stop quarks, we assume a charge survival probability of 38%, as discussed above, and exclude masses below 285 GeV. If charge flipping does not occur, we would obtain a higher mass limit, as indicated in Fig. 2c.

As seen in Figure 2, the observed limit exceeds the expected limit at various mass points by as much as 2.5 standard deviations, for all signals tested, due to the presence of the same few data events with high BDT discriminant values. This discrepancy reflects the excesses of data compared to background observed in Fig. 1 for the distributions both in beta (around 0.6) and dE/dx (around 2.8). We have compared these events with background events and conclude that the observed excess is consistent with a statistical fluctuation in the number of such events.

In summary, we have performed a search for charged, massive long-lived particles using 5.2 fb<sup>-1</sup> of integrated luminosity with the D0 detector. We find no evidence of signal and set 95% C.L. cross-section upper limits of order 0.01 pb for pair-produced CMLLPs of mass 200-300 GeV. At 95% C.L. we exclude pair-produced long-lived stop quarks with mass below 285 GeV, gaugino-like charginos below 267 GeV, and higgsino-like charginos below 217 GeV. These are presently the most restrictive limits for chargino CMLLPs, with about a factor of five improvement over the previous D0 cross section limits [10].

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- M. Byrne, C. Kolda, and P. Regan, Phys. Rev. D 66, 075007 (2002).
- [2] K. Kohri et al., Phys. Lett. B 682, 337 (2010).
- [3] K. Nakamura *et al.* (Particle Data Group), J. Phys. G 37, 075021 (2010). see Big Bang Nucleosynthesis (rev.), Section 20.5 and references cited therein.
- [4] J. Gunion and S. Mrenna, Phys. Rev. D 62, 015002 (2000).
- [5] C. Chen, M. Drees, and J. Gunion, Phys. Rev. D 55, 330 (1997).
- [6] G. F. Giudice and A. Rattazzi, Phys. Rep. 322, 419 (1999).
- [7] M. Strassler, arXiv:hep-ph/0607160.
- [8] M. Strassler and K. Zurek, Phys. Lett. B 651, 374, (2007).
- [9] D0 Collaboration, V. M. Abazov et al., Nucl. Instrum. Methods Phys. Res. A 565, 463 (2006); D0 Collaboration, M. We-

ber, Nucl. Instrum. Methods Phys. Res. A **566**, 182 (2006); D0 Collaboration, M. Abolins *et al.*, Nucl. Instrum. Methods Phys. Res. A **584**, 75 (2007).

- [10] D0 Collaboration, V. M. Abazov et al., Phys. Rev. Lett. 102, 161802, (2009).
- [11] CDF Collaboration, D. Acosta *et al.*, Phys. Rev. Lett. **90**, 131801 (2003).
- [12] CDF Collaboration, T. Aaltonen *et al.*, Phys. Rev. Lett. **103**, 021802 (2009).
- [13] ALEPH, DELPHI, L3 and OPAL Collaborations, notes LEPSUSYWG/02-05.1 and LEPSUSYWG/02-09.2.
- [14] CMS Collaboration, V. Khachatryan *et al.*, J. High Energy Phys. 03, 024 (2011).
- [15] ATLAS Collaboration, G. Aad et al., Phys. Lett. B 698, 353 (2011).
- [16] The D0 coordinate system is cylindrical with the *z*-axis along the proton beam direction, and the polar and azimuthal angles are denoted by  $\theta$  and  $\phi$ , respectively. The pseudorapidity is defined as  $\eta = -\ln[\tan(\theta/2)]$ .
- [17] The transverse mass is defined by  $M_T = \sqrt{(E_T + \not\!\!E_T)^2 (p_x + \not\!\!E_x)^2 (p_y + \not\!\!E_y)^2}$ , where  $E_T$  is the

- [18] The requirement  $M_T < 200$  GeV is customarily used to select W events. There is negligible contamination of potential signal events in this data sample.
- [19] T. Sjostrand et al., Comput. Phys. Commun. 135, 238 (2001).
- [20] The code for generating stop quarks is at http://projects.hepforge.org/pythia6/examples/main78.f
- [21] R. Brun and F. Carminati, CERN Program Library Long Writeup W5013, 1993 (unpublished).
- [22] M. Fairbairn et al., Phys. Rept. 438, 1 (2007).
- [23] R. Mackeprang and A. Rizzi, Eur. Phys. J. C50, 353 (2007).
- [24] R. Mackeprang and D. Milstead, Eur. Phys. J. C66, 493 (2010).
- [25] R. Mackeprang, CERN-THESIS-2007-109, page 43.
- [26] A. Hoecker et al., PoS ACAT 040 (2007),
- arXiv:physics/0703039.
- [27] T. Andeen *et al.*, FERMILAB-TM-2365 (2007).
- [28] W. Fisher, FERMILAB Report No. TM-2386-E, (2007).
- [29] W. Beenakker et al., Phys. Rev. Lett. 83, 3780 (1999).