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# Neutrino Oscillation measurements and BSM physics searches with Neutrino Telescopes in the Mediterranean

Paschal Coyle

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Submitted on 5 Mar 2020

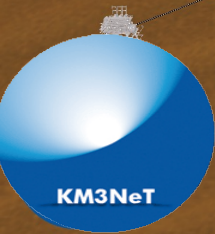
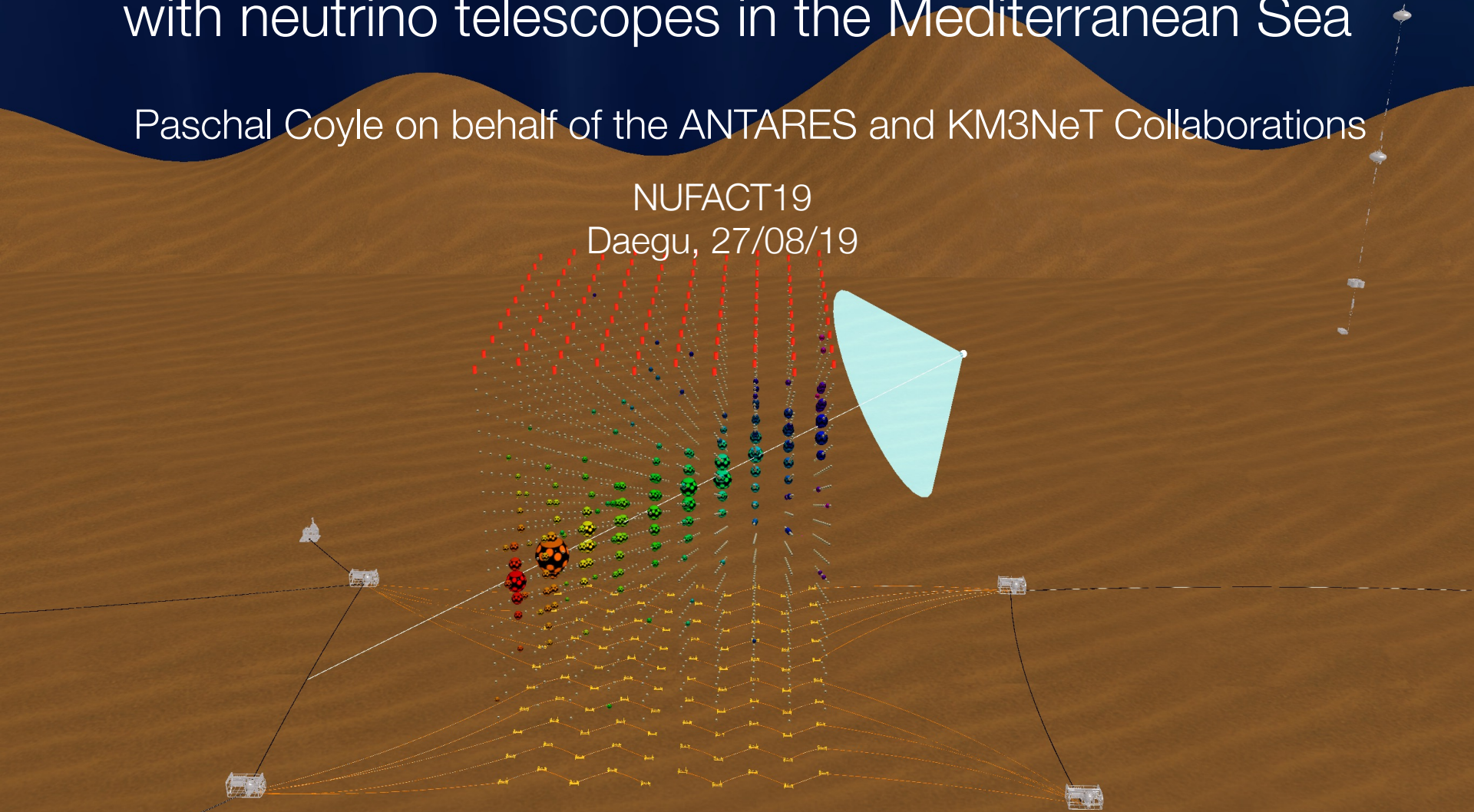
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# Oscillation measurements and BSM physics searches with neutrino telescopes in the Mediterranean Sea

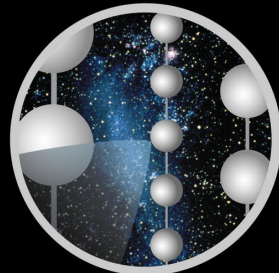
Paschal Coyle on behalf of the ANTARES and KM3NeT Collaborations

NUFACT19  
Daegu, 27/08/19



# Large Volume Neutrino Telescopes

ANTARES & KM3NeT

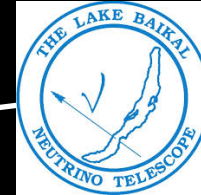


ICECUBE

IceCube



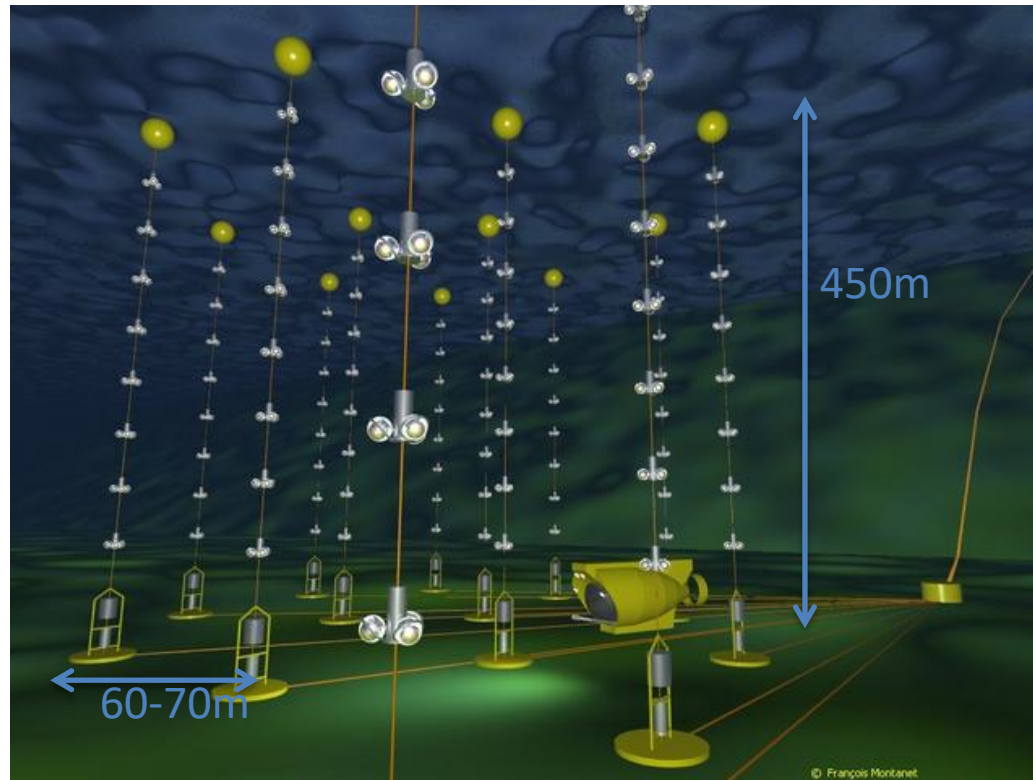
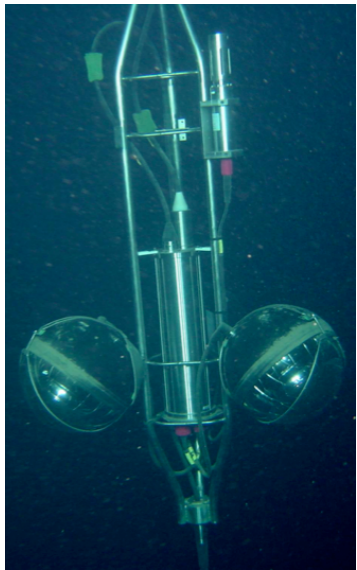
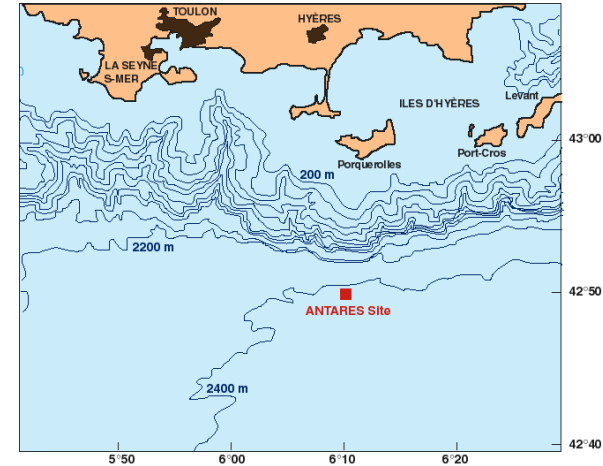
BAIKAL & GVD





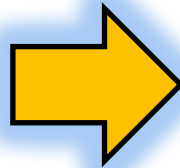
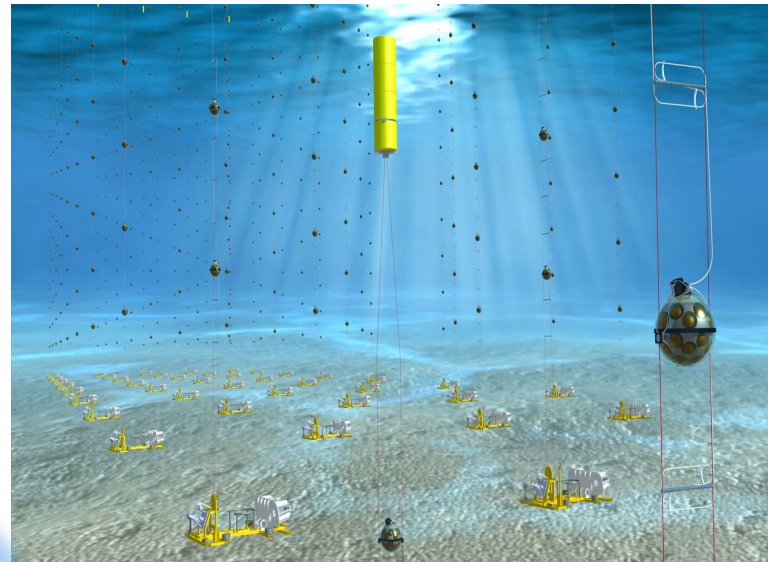
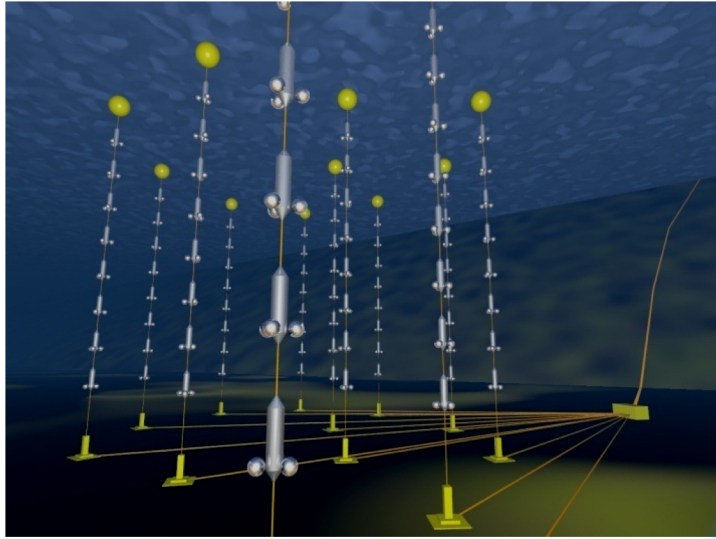
# ANTARES

- 42km offshore Toulon, depth 2475m
- Main Electro-Optic Cable/Junction Box 2001-2002
- Completed 2008
- 12 lines, ~70m spacing
- 25 storeys per line, 15m spacing
- 3x10-inch PMTs per storey
- Decommissioning 2017



12 lines, 900 OMs

3 Building Blocks (3\*115 lines, ~3\*2000 OMs)

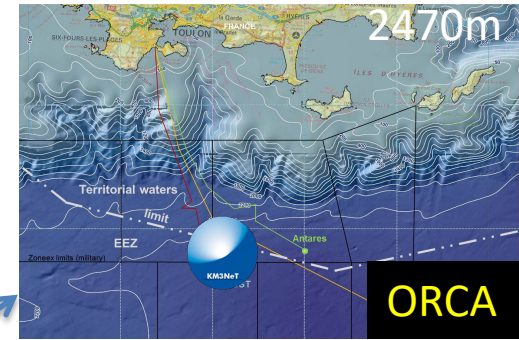


- 31 x 3" PMTs
- Uniform angular coverage
- Directional information
- Digital photon counting
- Reduced ageing
- All data to shore



# KM3NeT

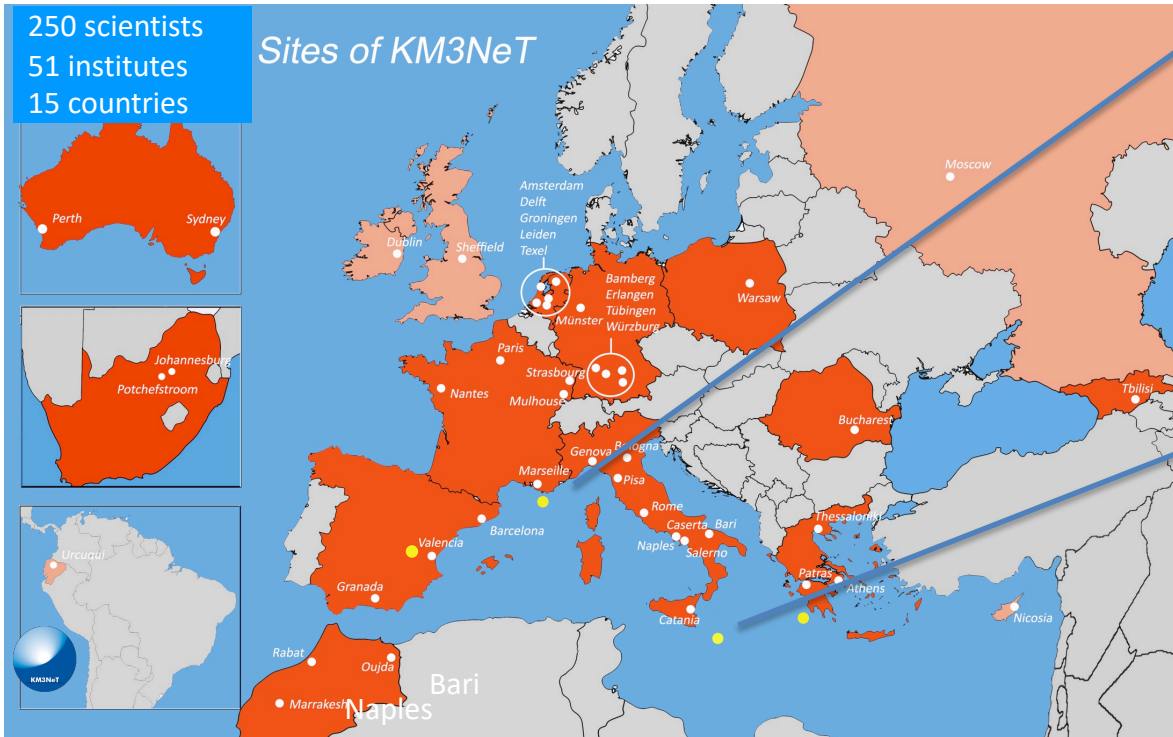
Multi-site, deep-sea infrastructure  
Selected by ESFRI roadmap  
Single collaboration, Single technology



Oscillation Research  
with Cosmics In the Abyss



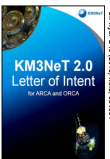
Astroparticle Research  
with Cosmics In the Abyss



[KM3NeT 2.0: Letter of Intent](http://dx.doi.org/10.1088/0954-3899/43/8/084001)


<http://dx.doi.org/10.1088/0954-3899/43/8/084001>

J. Phys. G: Nucl. Part. Phys. 43 (2016) 084001



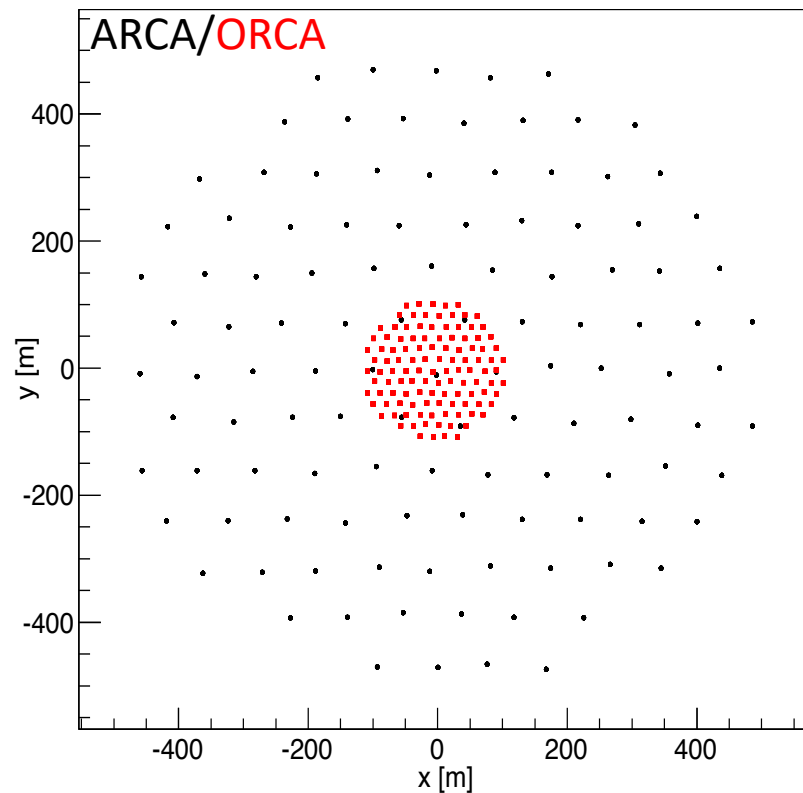
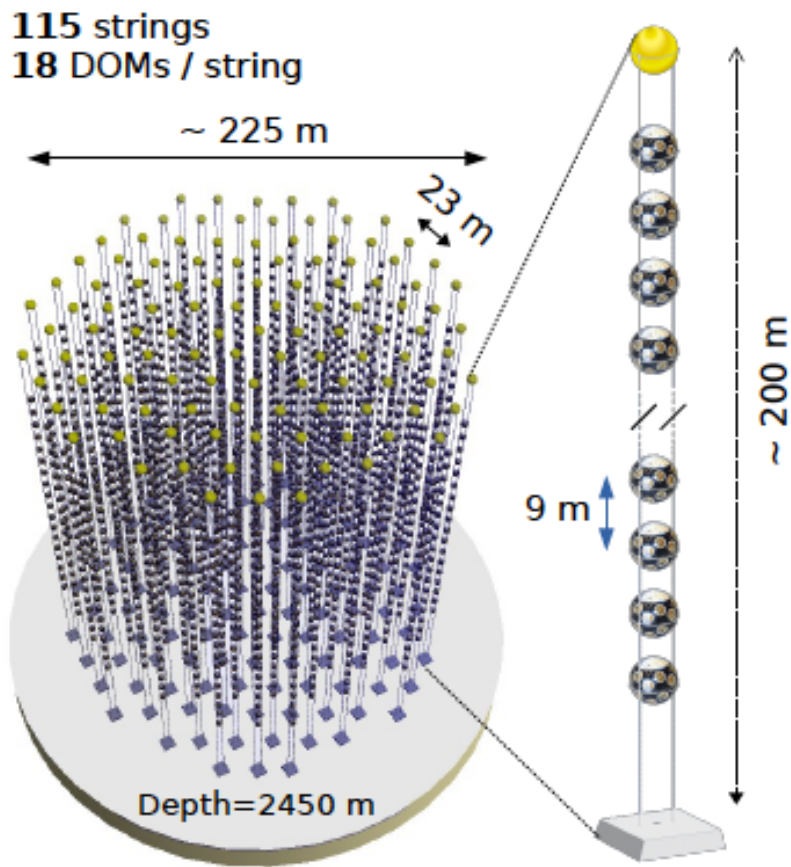
Connection nodes of

european  
 multidisciplinary  
 seafloor & water column  
 observatory





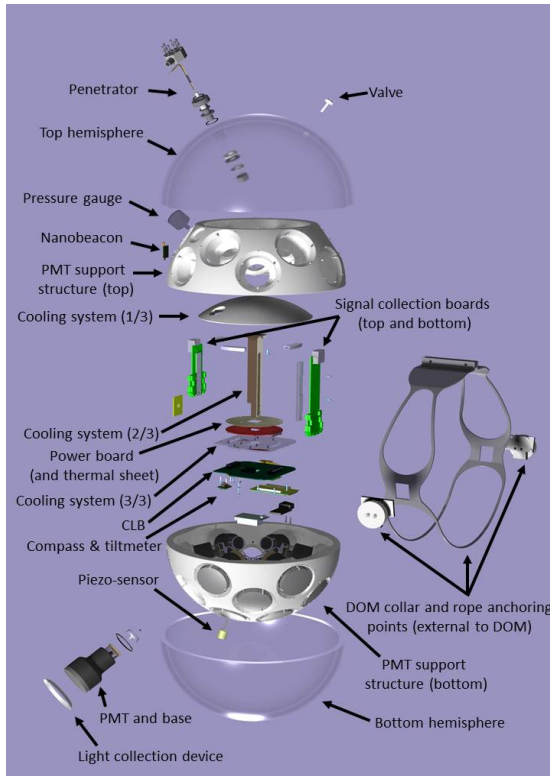
# The KM3NeT Building Block



- **31 PMTs / DOM**
- **Total: 64k\*3" PMTs**

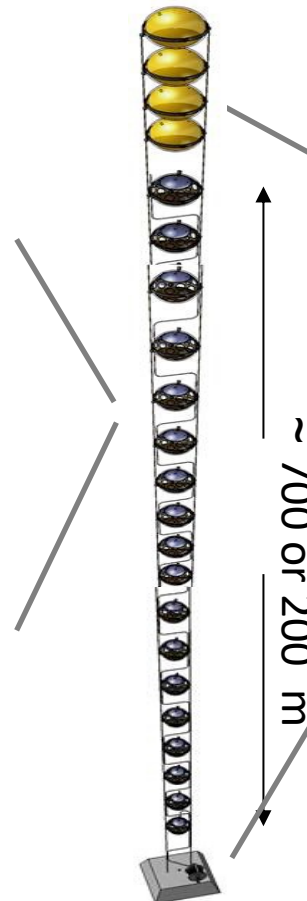
	ORCA	ARCA
String spacing	20 m	90 m
OM spacing	9 m	36 m
Instrumented mass	8 Mton	500*2 Mton

## Digital Optical Module



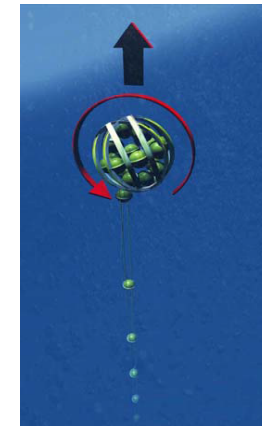
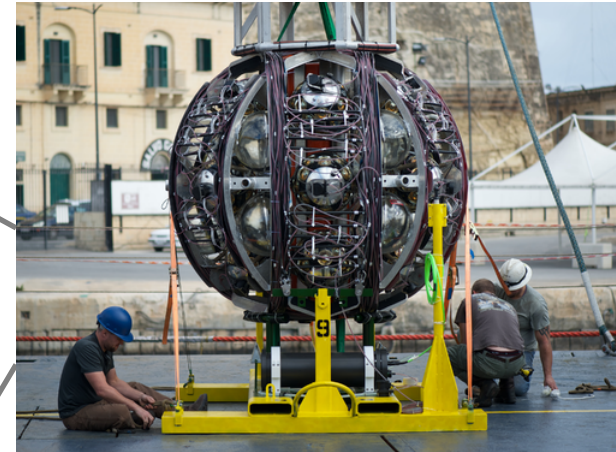
- 31 x 3" PMTs
- Gbit/s on optical fibre
- Hybrid White Rabbit
- LED flasher & acoustic piezo
- Tiltmeter/compass

## Detection unit



- 2 dyneema ropes
- Oil filled PVC tube
- Low drag
- Low cost

## Deployment Vehicle



- Rapid deployment
- Multiple strings/sea campaign
- Autonomous/ROV unfurling
- Reuseable



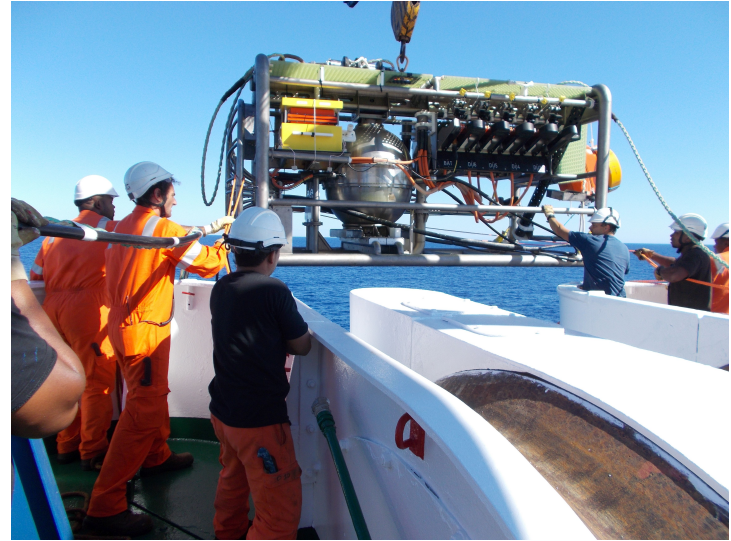


# ORCA: Some construction milestones

Main Cable: dec 2015, sept 2018



Node 1: sept 2016, sept 2018

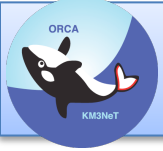


DUs: feb(1), may(1), july(2) 2019



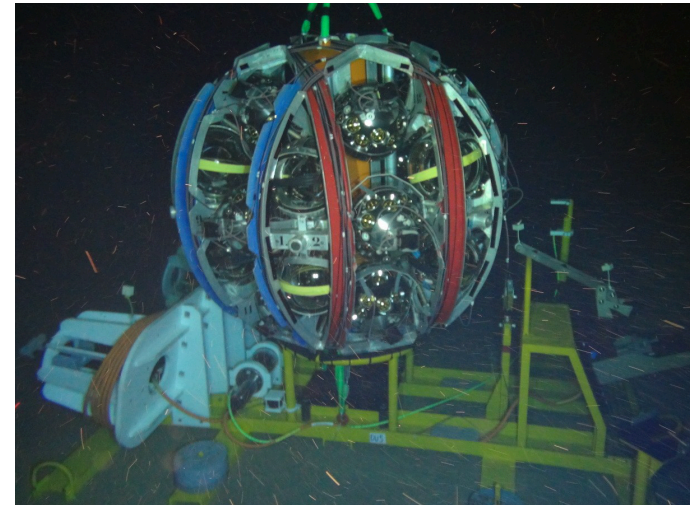
Instrumentation module: may 2019



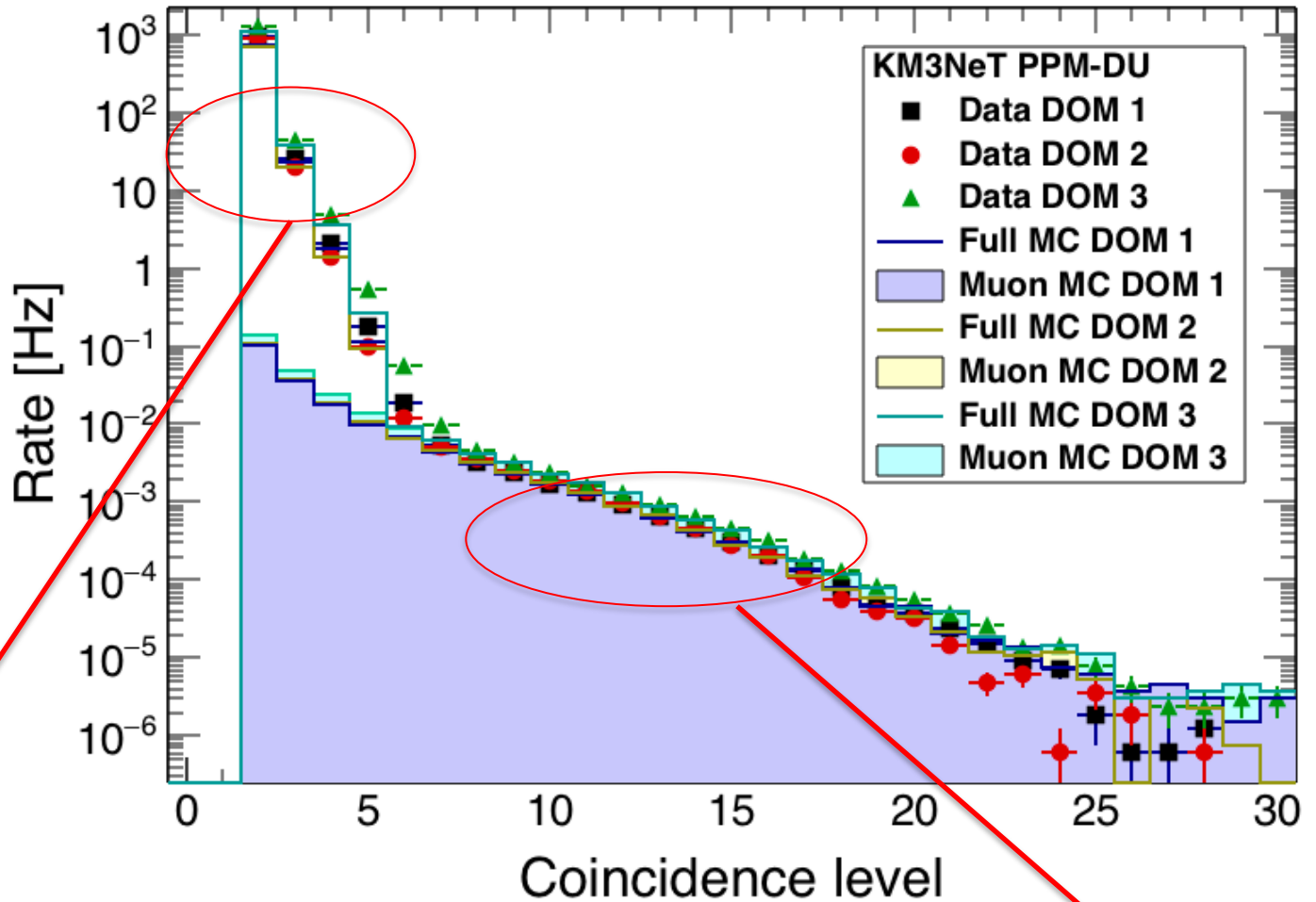


# detection unit deployment/connection

<https://youtu.be/dMjN93H7Nvo>



# DOM coincidences



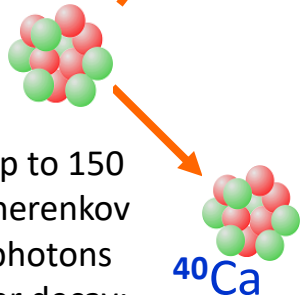
$^{40}\text{K}$

Atmospheric muons

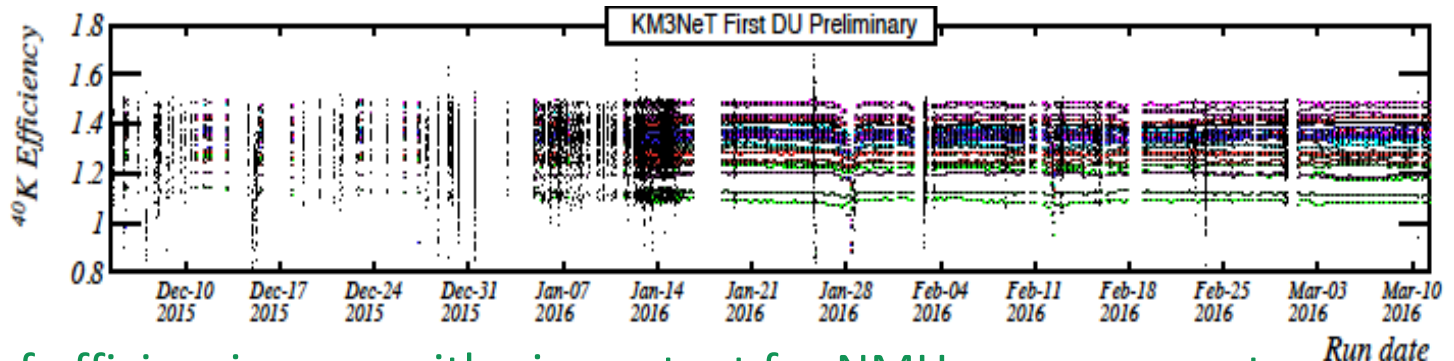
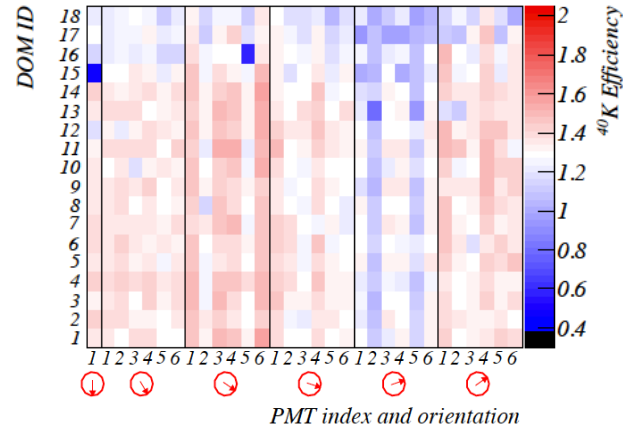
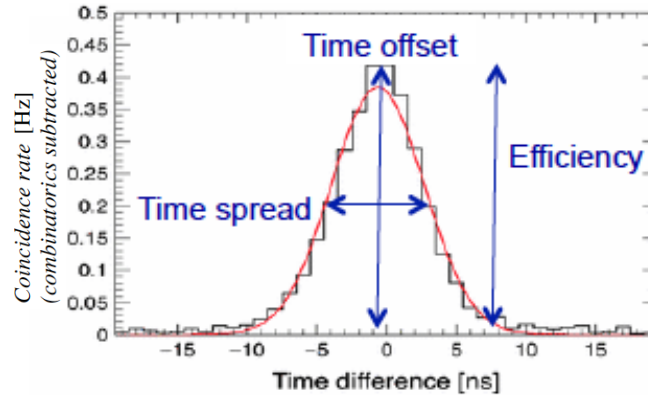
# $^{40}\text{K}$ : Inter-PMT Calibration



$^{40}\text{K}$   $\rightarrow$   $e^-$  ( $\beta$  decay)



Up to 150 Cherenkov photons per decay; stable  $^{40}\text{K}$  concentration

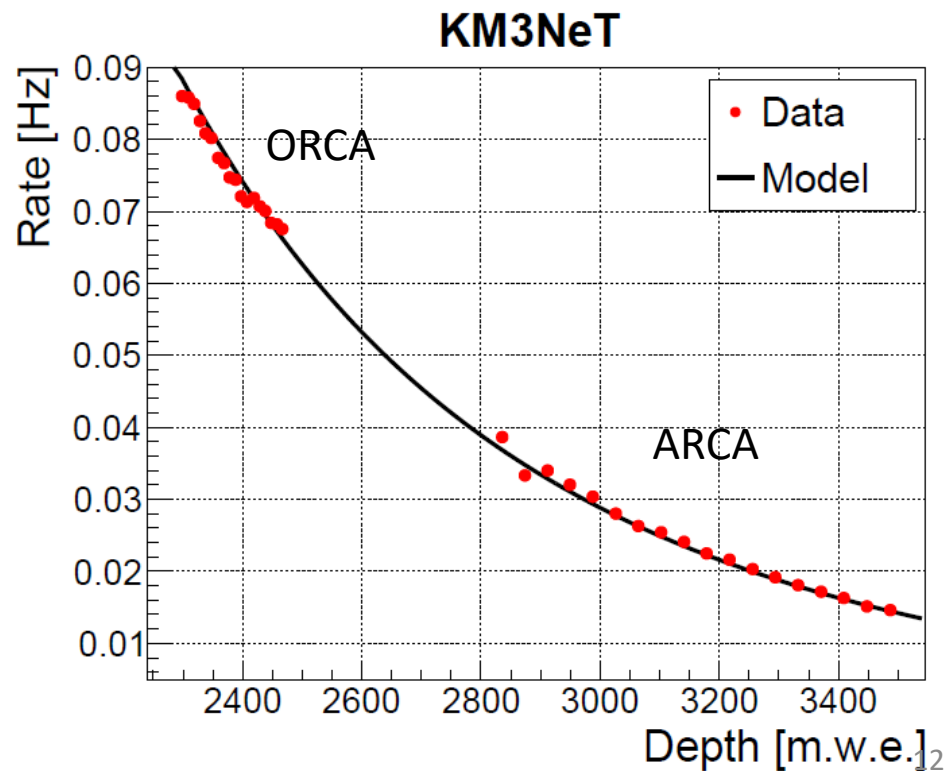
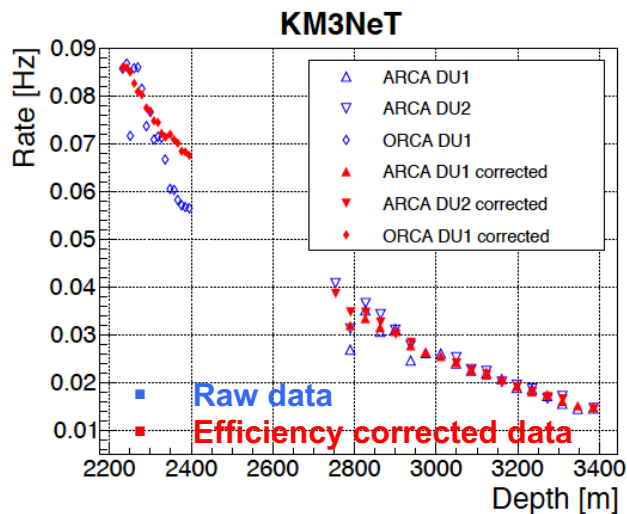
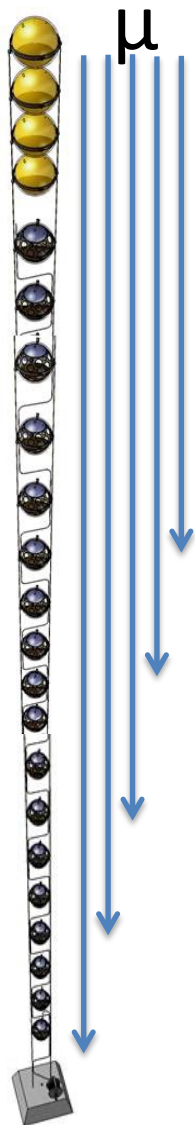


Knowledge of efficiencies vs zenith - important for NMH measurement

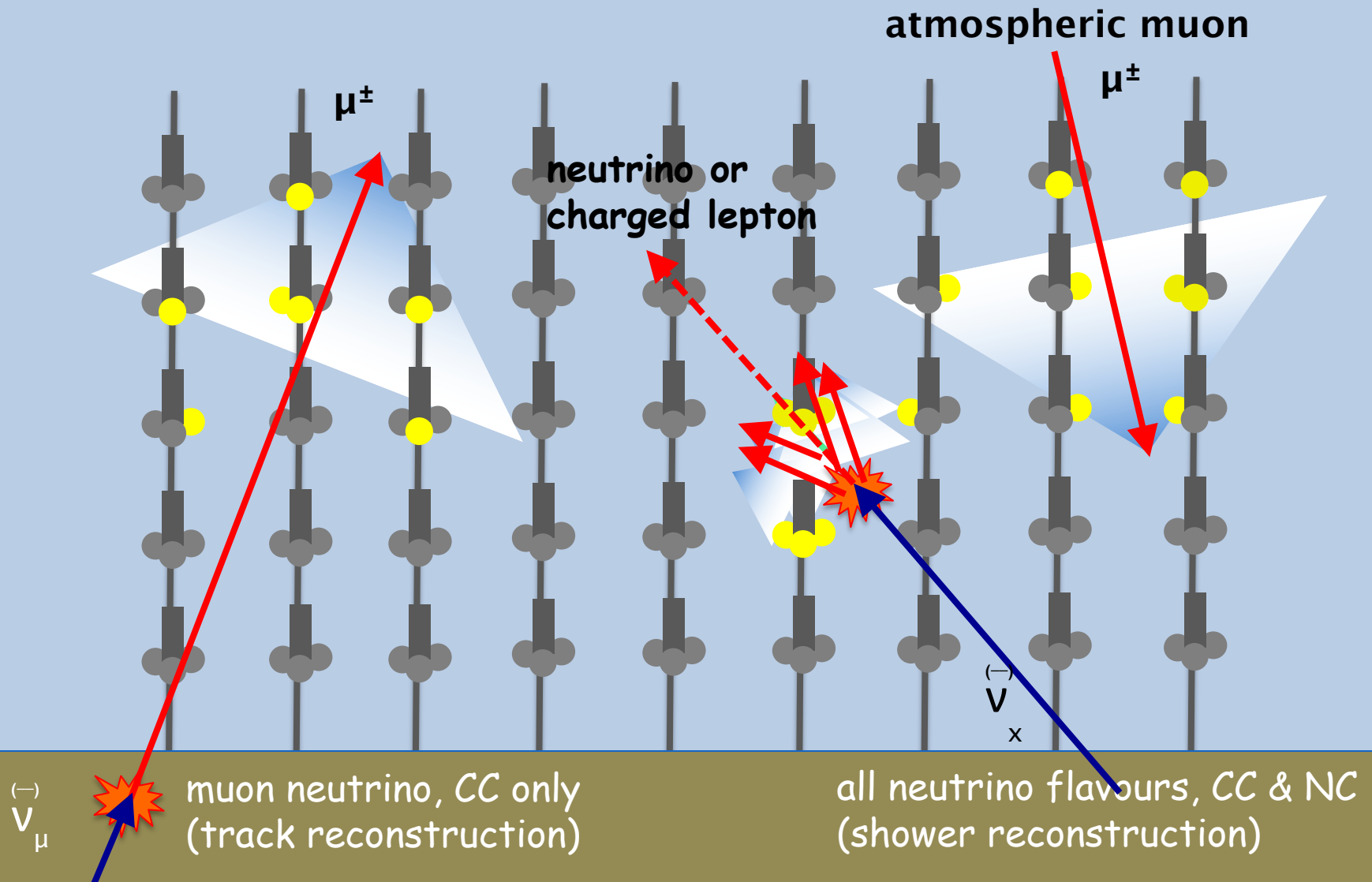


# ARCA + ORCA Muon Depth Dependence

Joint ARCA/ORCA analysis measures the muon flux attenuation over > 1 km length  
e-Print: arXiv:1906.02704

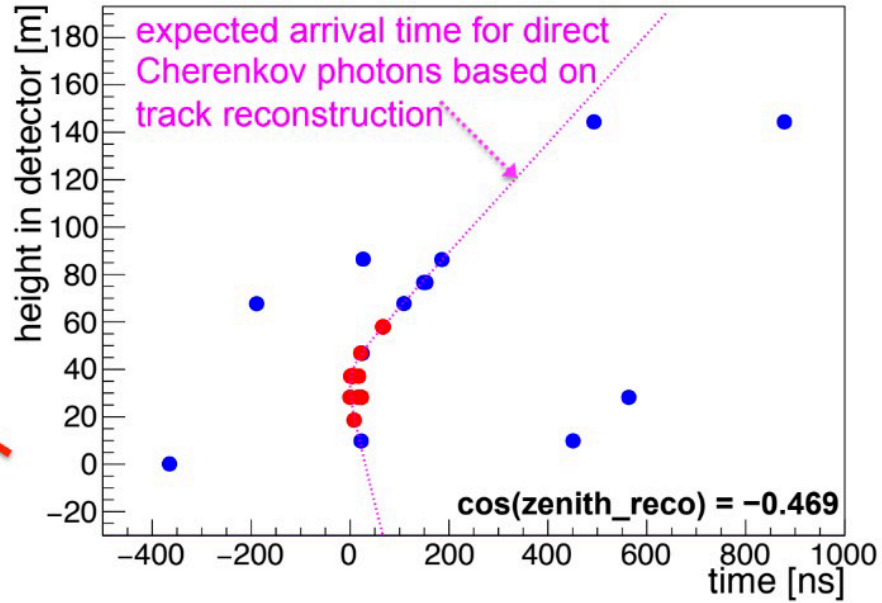


# Neutrino signatures

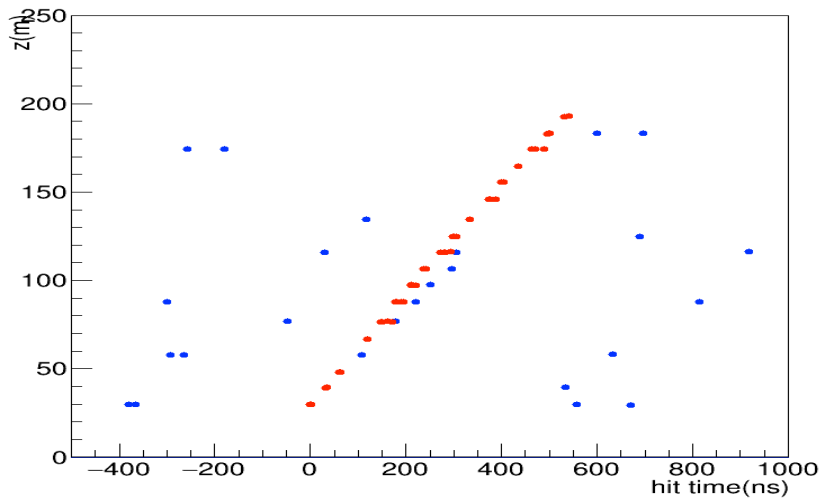
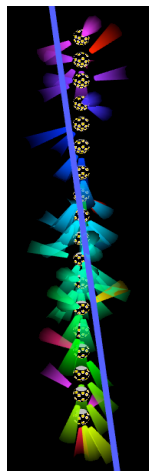


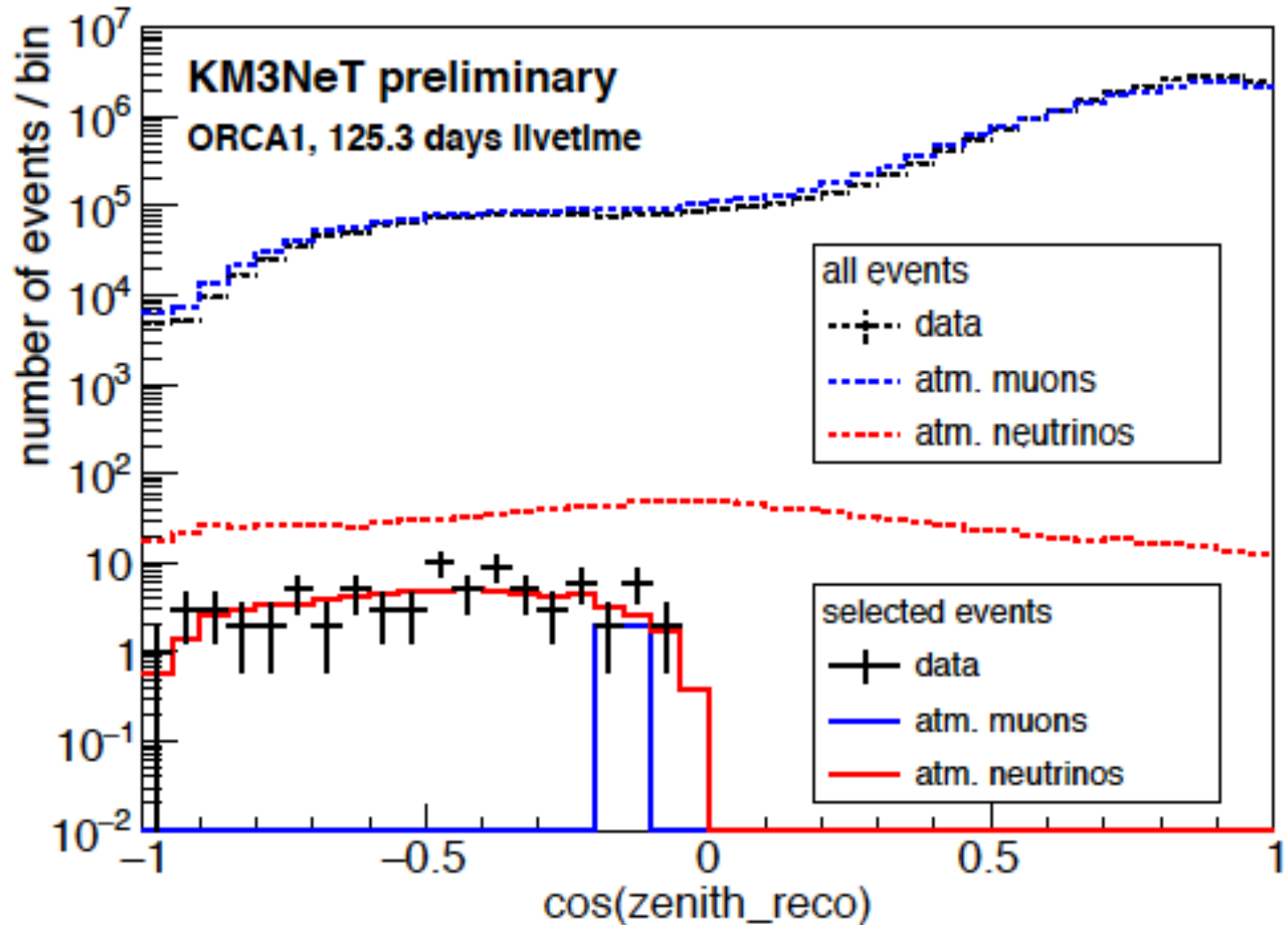
# ORCA1: neutrinos

event=1668, run=2974, #hits=26, cos(zenith\_reco)=-0.469  
DU 2



Evt: id=3860 run\_id=2609 #hits=87 #mc\_hits=0 #trks=0 #mc\_trks=0



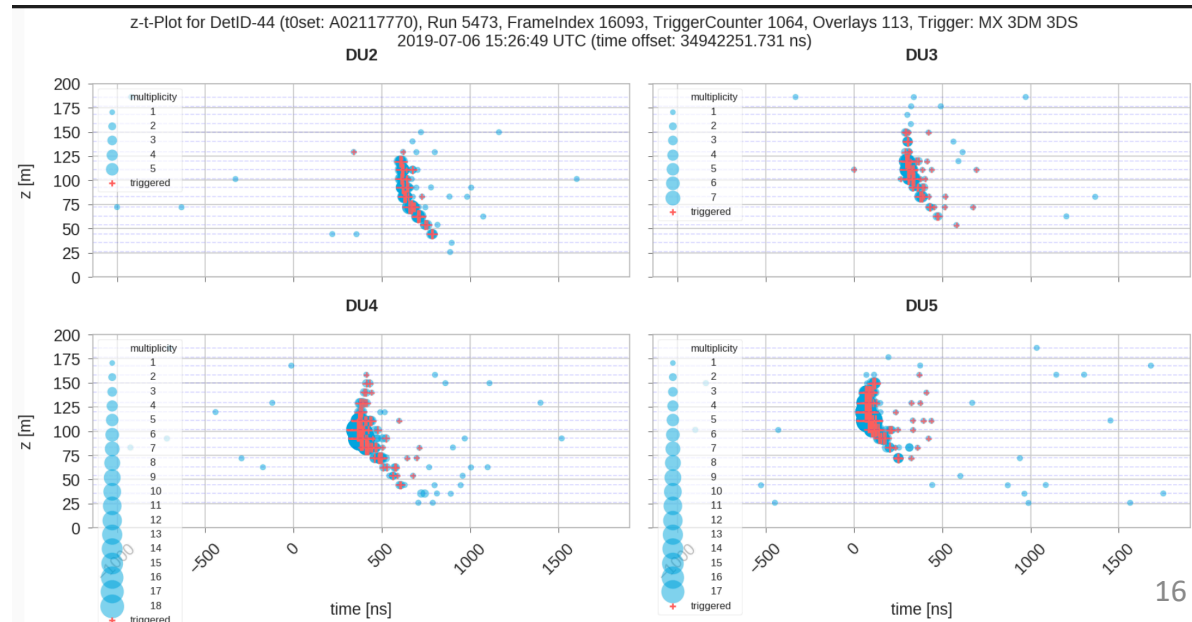
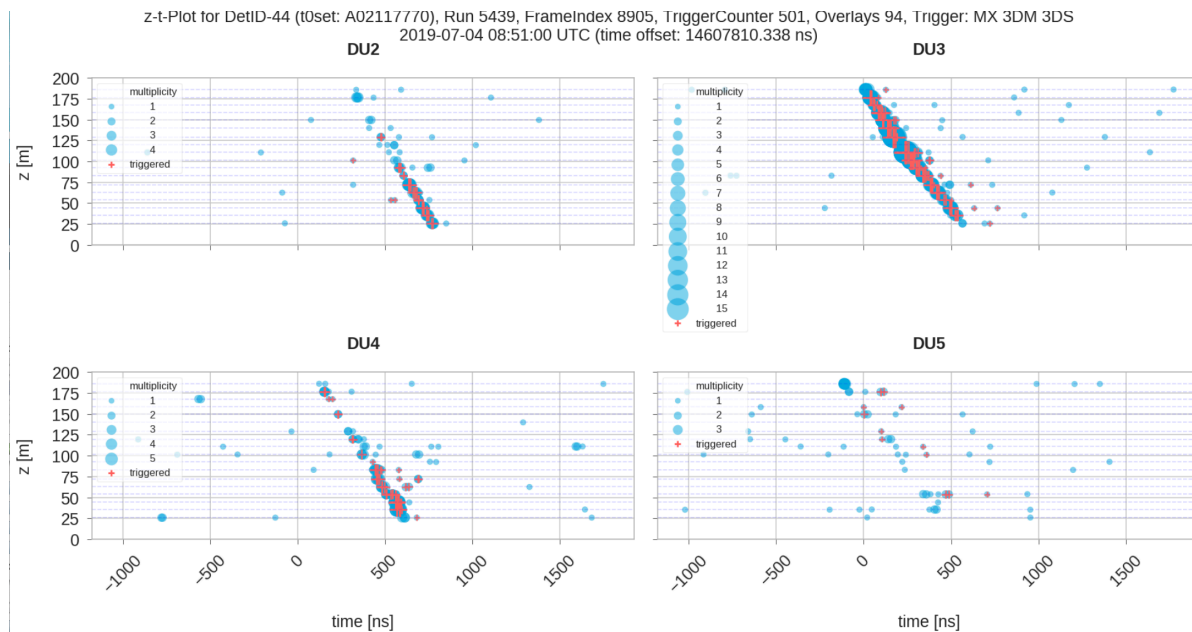
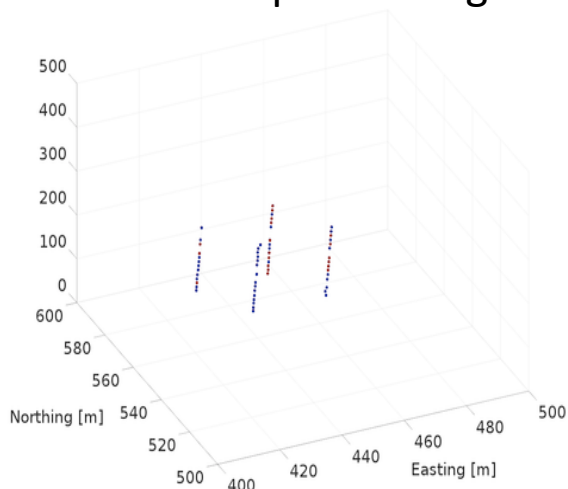






# ORCA4: atmospheric muons

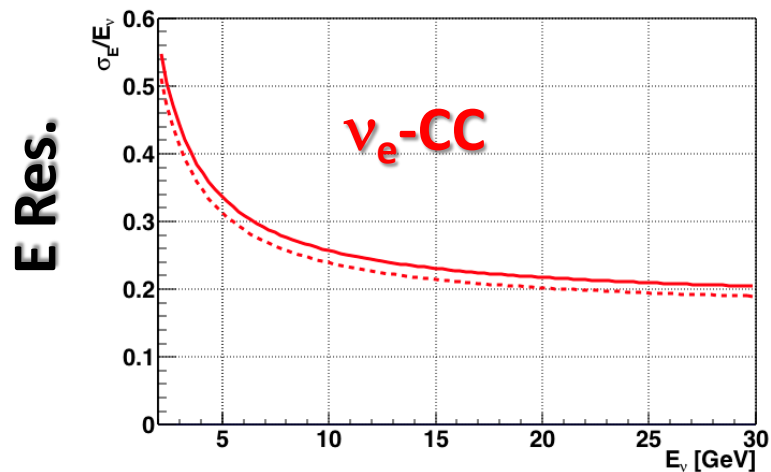
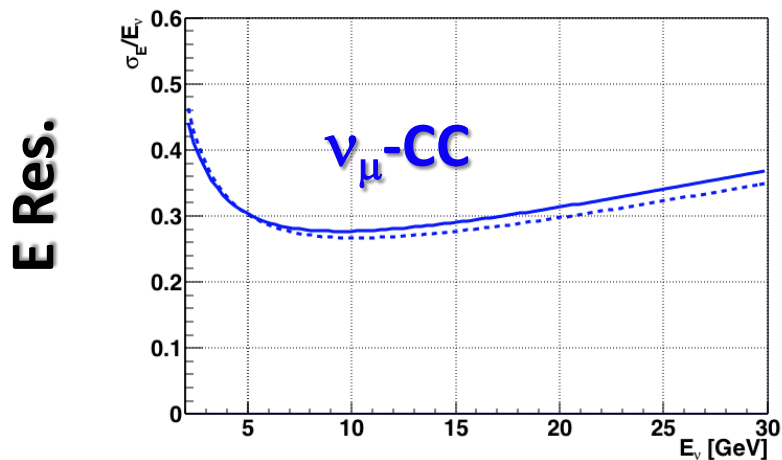
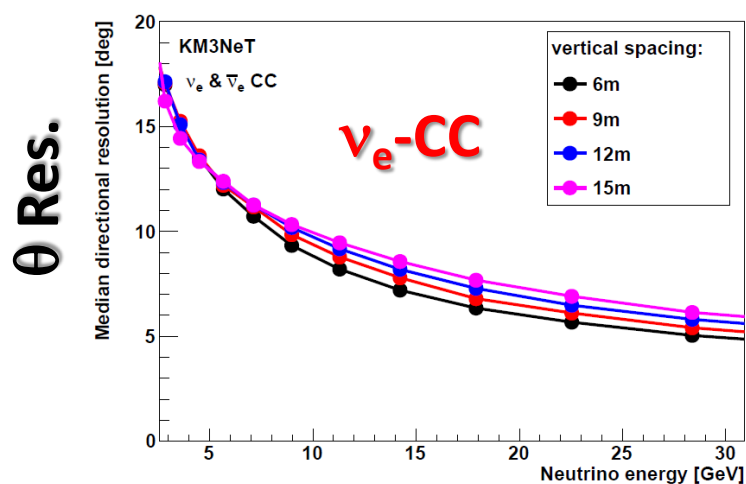
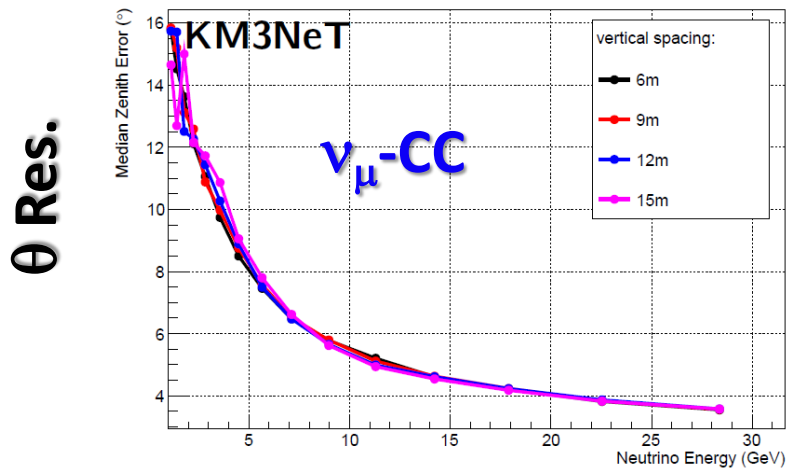
## Acoustic positioning





# ORCA: reconstruction performance

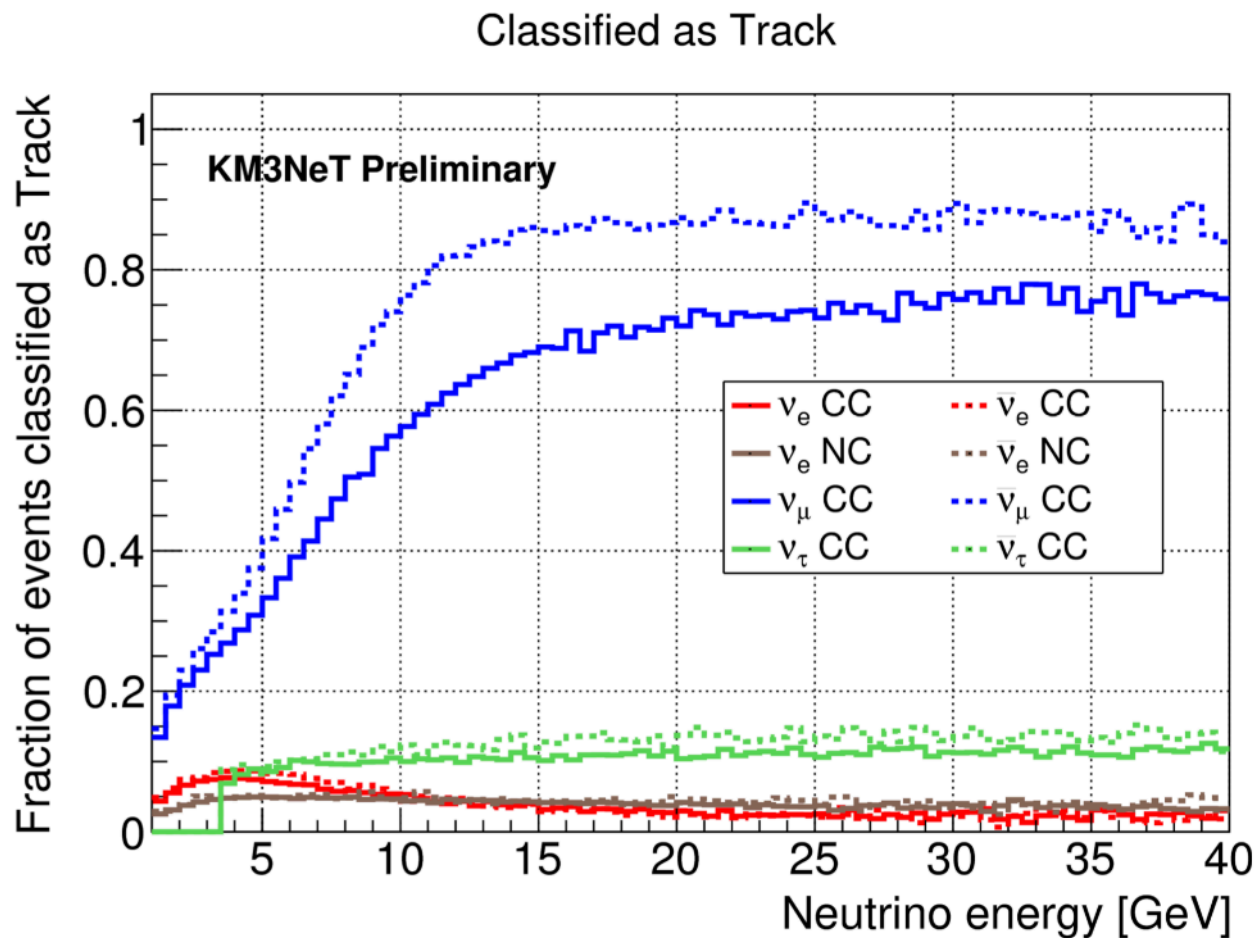
- Angular resolution: Better than 10 degrees at relevant energies
- Energy resolution:  $\sim 25\%$  (Close to intrinsic limit [arXiv:1612.05621](https://arxiv.org/abs/1612.05621))





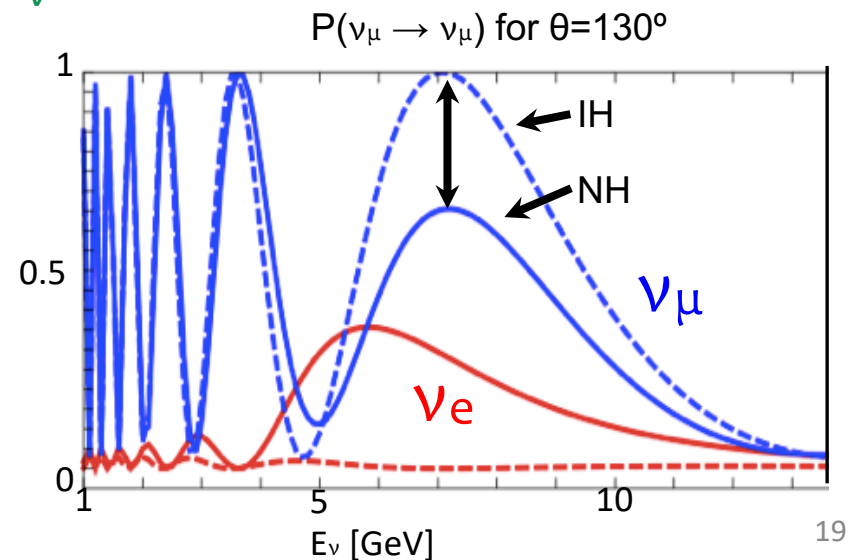
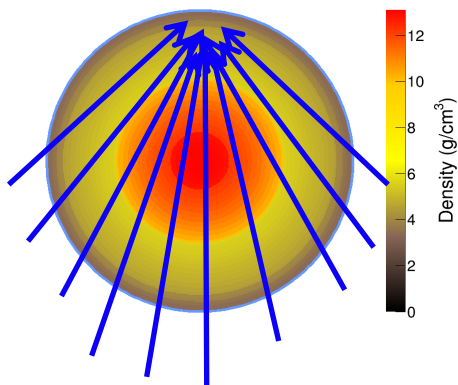
# ORCA: shower/track identification

Discrimination of track-like and shower-like events  
via Random Decision Forest



# Oscillations with atmospheric neutrinos

- A “free beam” of known composition ( $\nu_e, \nu_\mu$ )
- A “free cavern” of known/uniform composition
- **Wide range of baselines and energies**
- Oscillation pattern distorted by Earth matter effects  
maximum difference IH  $\square$  NH for resonance in  
Earth mantle:  $\theta=130^\circ$  (7645 km) and  $E_\nu = 7$  GeV





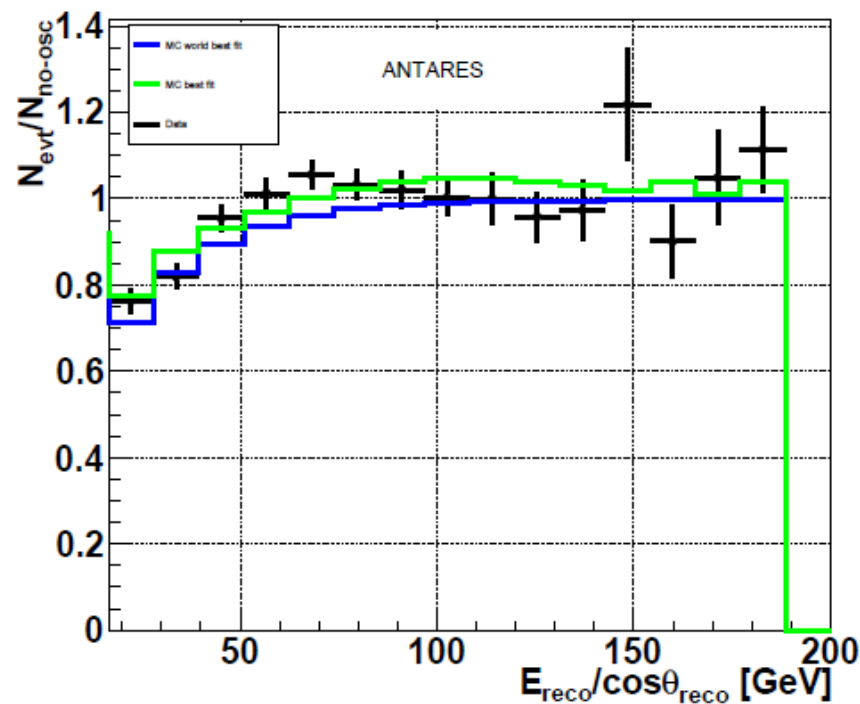
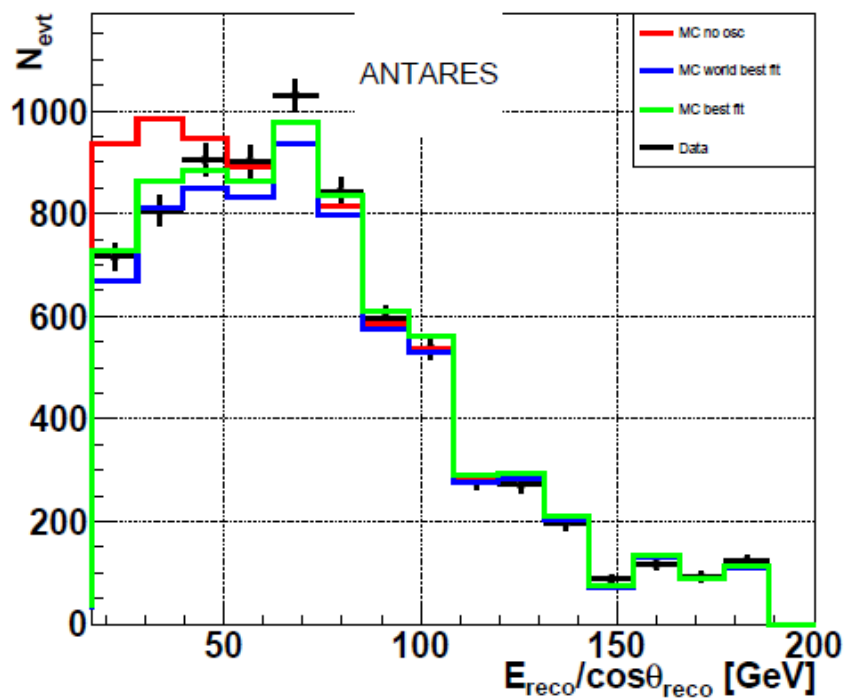
# ANTARES neutrino oscillations

arXiv:1812.08650v3 [hep-ex] 21 May 2019

Data sample: 9 years (2007-2016) -2830 days lifetime

7710 events selected: Tracks only

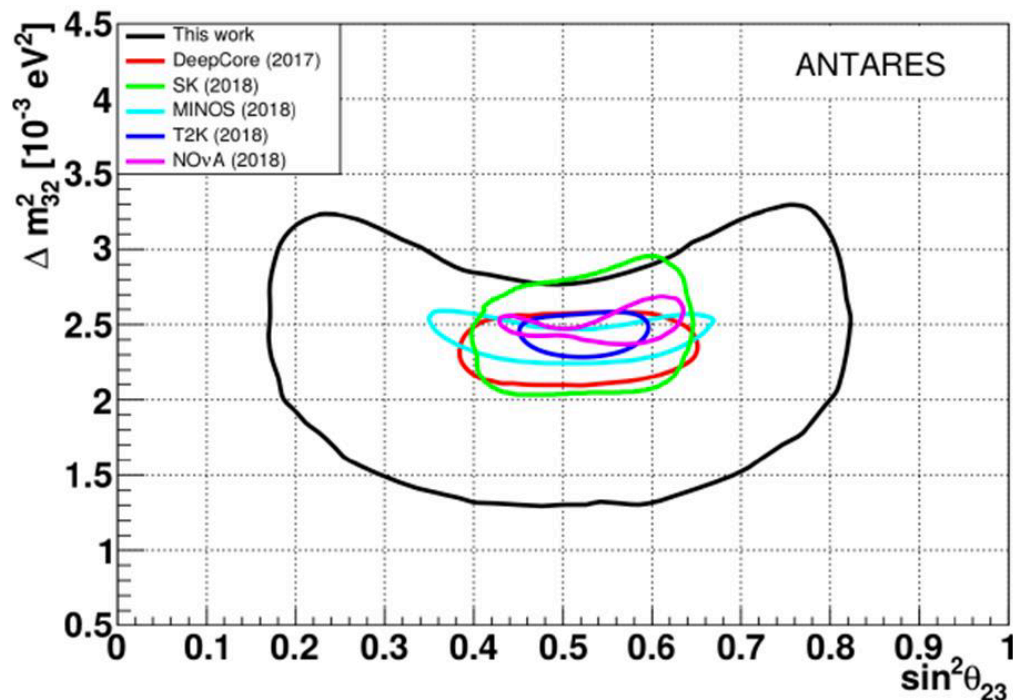
A binned likelihood fit is performed in two dimensions ( $E_{\text{reco}}$ ,  $\cos\theta_{\text{reco}}$ )





# ANTARES: oscillations parameters

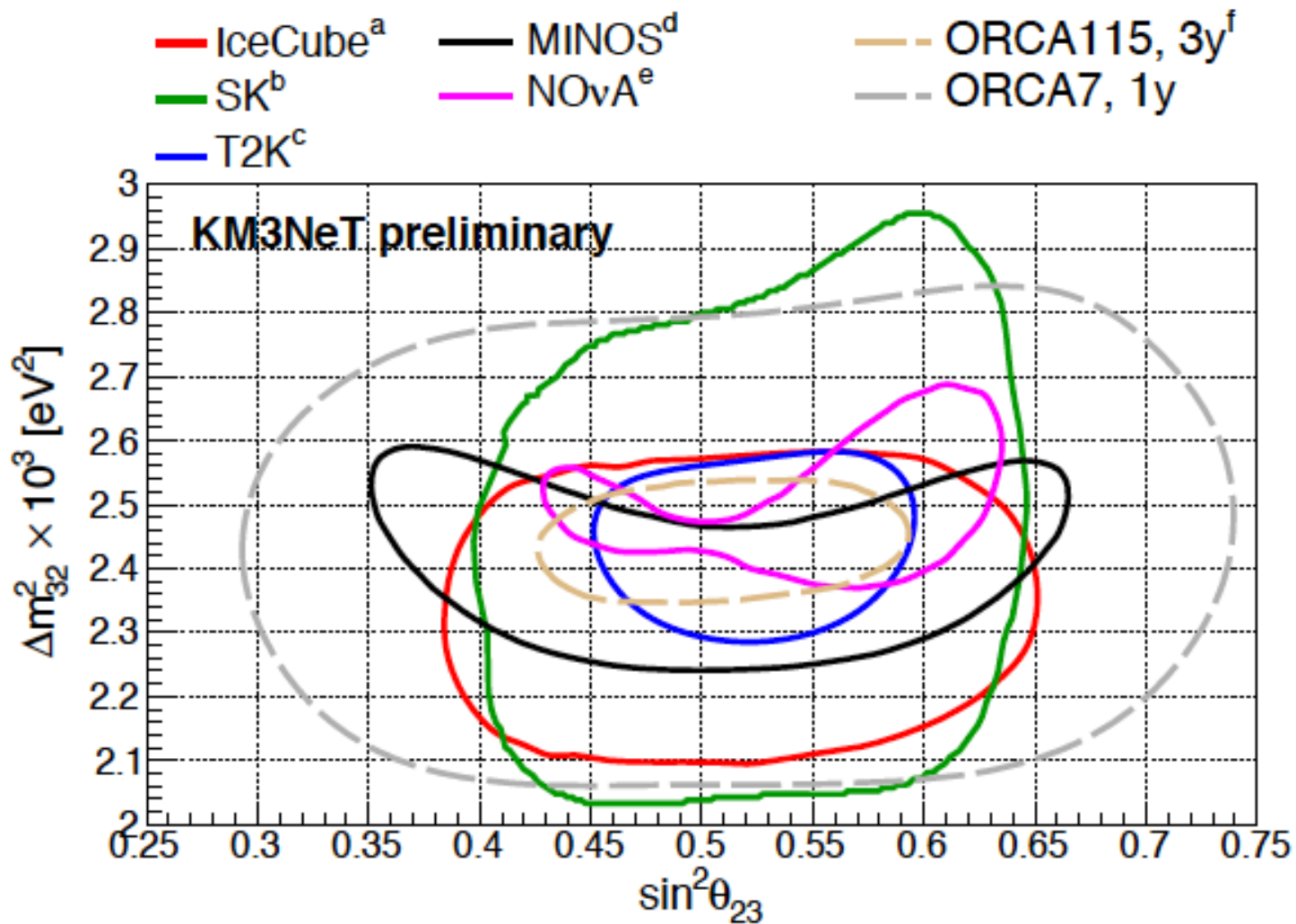
Parameter	Prior	Fit result
$\Delta m_{32}^2$ [ $10^{-3}$ eV $^2$ ]	none	$2.0^{+0.4}_{-0.3}$
$\theta_{23}$ [ $^\circ$ ]	none	$45^{+12}_{-11}$
$n_\nu$	none	$0.81^{+0.10}_{-0.09}$
$\nu/\bar{\nu}$ [ $\sigma$ ]	$0.0 \pm 1.0$	$1.10^{+0.64}_{-0.56}$
$\Delta\gamma$	$0.00 \pm 0.05$	$-0.003 \pm 0.036$
$N_\mu$	$740 \pm 120$	$414^{+48}_{-24}$
$\theta_{13}$ [ $^\circ$ ]	$8.41 \pm 0.28$	$8.41 \pm 0.28$
$M_A$ [ $\sigma$ ]	$0.0 \pm 1.0$	$0.0 \pm 1.0$



The non-oscillation hypothesis rejected at 4.6 sigma

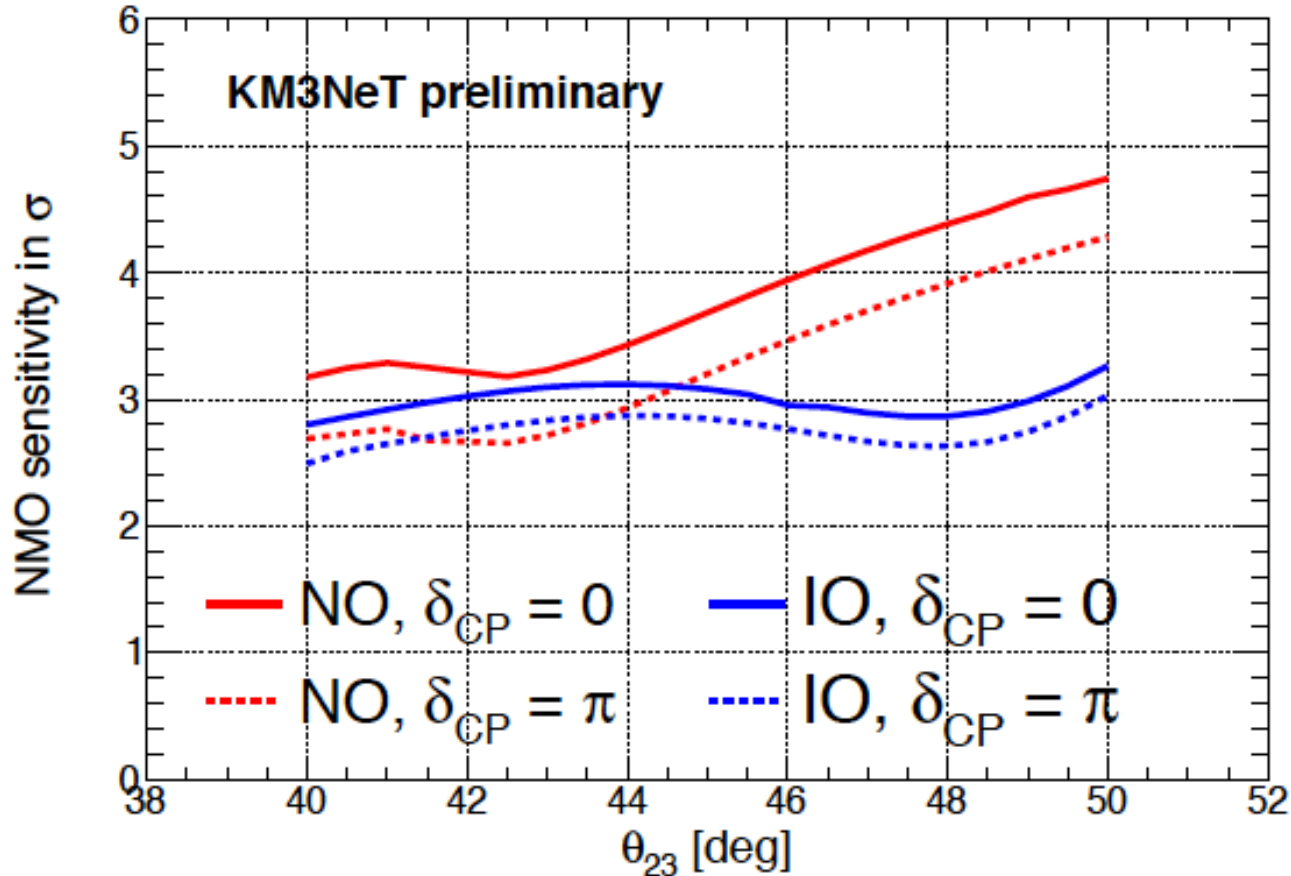


# KM3NeT/ORCA: oscillation parameters





# Sensitivity to neutrino mass hierarchy



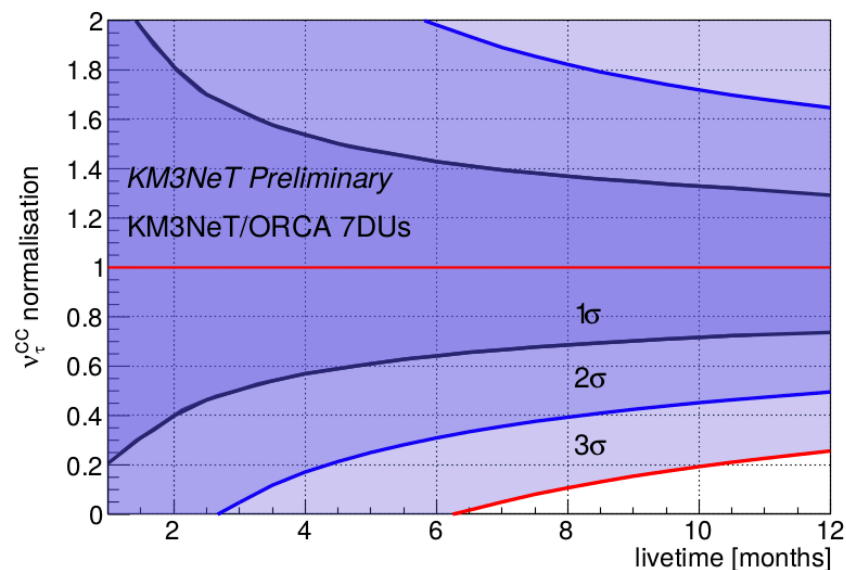
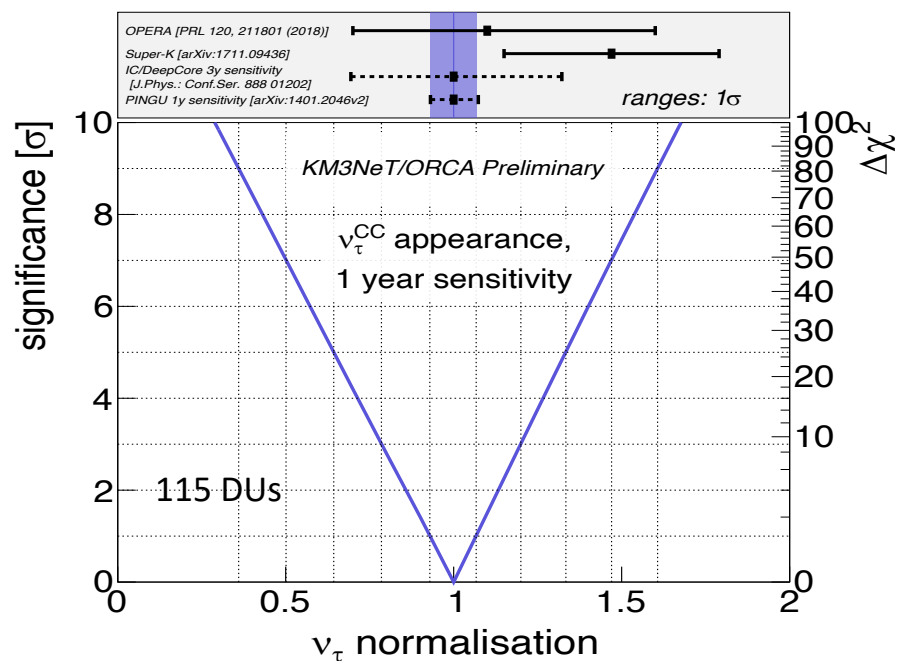
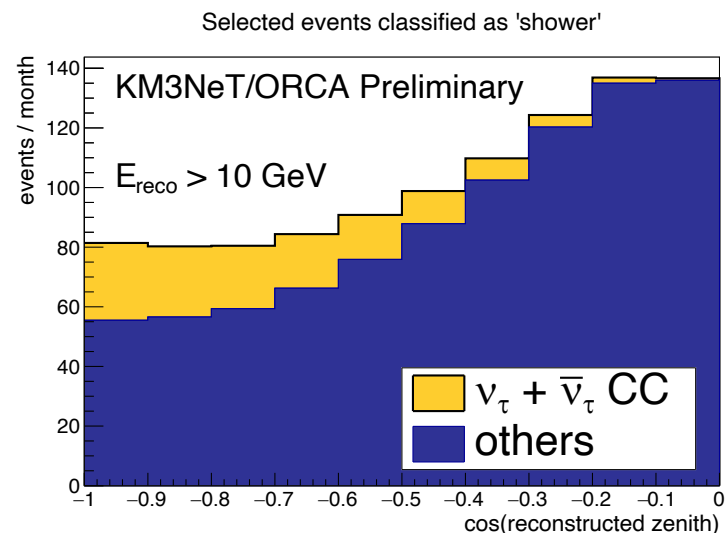
- $\sim 3\sigma$  MH sensitivity in 3 years
- The combination of NH and upper octant of  $\theta_{23}$  gives improved sensitivity
- The value of  $\delta_{cp}$  has small but non-negligible impact on sensitivity





# KM3NeT/ORCA: Tau neutrino appearance

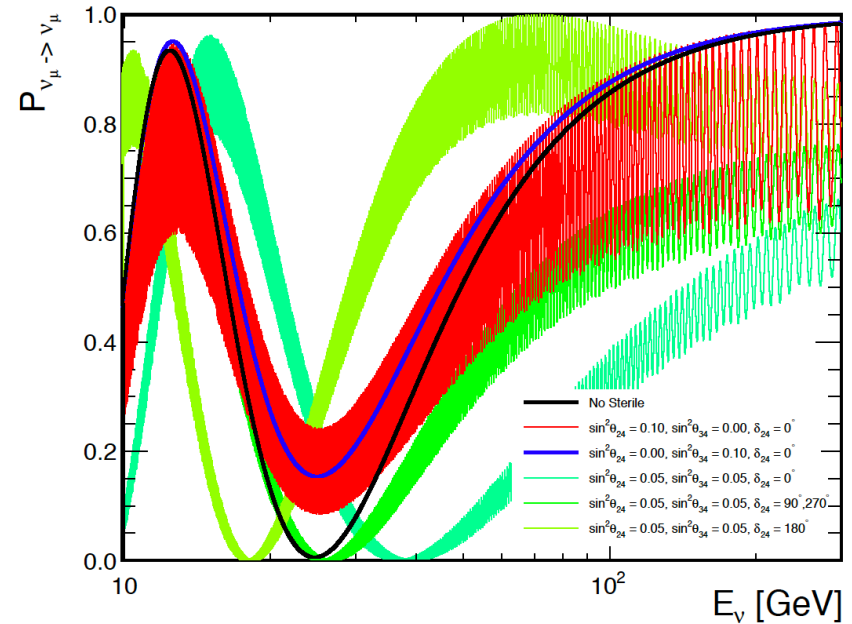
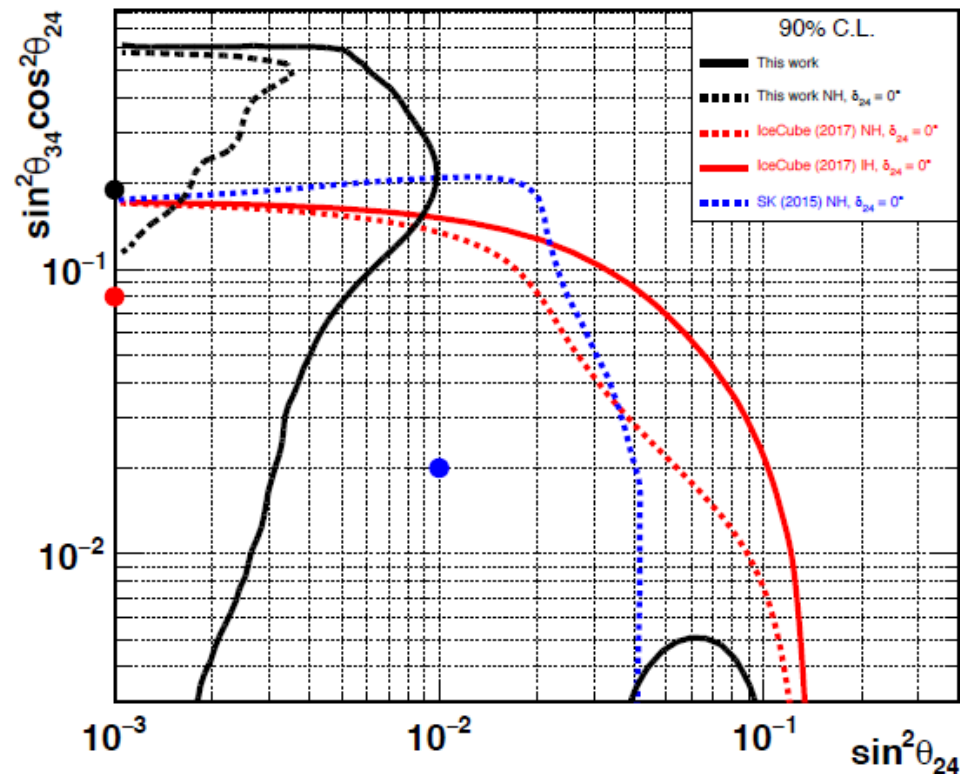
- $\nu_\tau$  appearance tests PMNS unitarity and BSM theories
- 30% deviations allowed by world data
- $\approx 3k$   $\nu_\tau$  CC events/year with full ORCA
- Rate constrained within  $\approx 5$  (25)% for 115 (7) DUs in 1 year





# ANTARES: sterile neutrino (3+1)

Presence of sterile neutrino modify significantly the oscillation pattern



Parameter	Prior	Fit NH	Fit IH
$\theta_{24}$ [°]	none	$1.5^{+2.0}_{-5.0}$	$1.5^{+2.0}_{-5.0}$
$\theta_{34}$ [°]	none	$25.9^{+5.1}_{-4.2}$	$25.9^{+5.1}_{-4.2}$
$\delta_{24}$ [°]	none	$180 \pm 71$	$0 \pm 72$
$n_\nu$	none	$0.84^{+0.10}_{-0.09}$	$0.84^{+0.10}_{-0.09}$
$\nu/\bar{\nu}$ [σ]	$0.0 \pm 1.0$	$1.07^{+0.63}_{-0.55}$	$1.07^{+0.63}_{-0.55}$
$\Delta\gamma$	$0.00 \pm 0.05$	$-0.011 \pm 0.036$	$-0.011 \pm 0.036$
$\Delta m_{32}^2$ [ $10^{-3}$ eV <sup>2</sup> ]	none	$3.0^{+0.8}_{-0.6}$	$-3.0^{+0.6}_{-0.8}$
$\theta_{23}$ [°]	none	$52 \pm 8$	$52 \pm 8$
$\theta_{13}$ [°]	$8.41 \pm 0.28$	$8.41 \pm 0.28$	$8.41 \pm 0.28$
$M_A$ [σ]	$0.0 \pm 1.0$	$0.11^{+0.93}_{-0.97}$	$0.11^{+0.93}_{-0.97}$

$$|U_{\mu 4}|^2 < 0.007 \text{ (0.13) at 90\% (99\%) CL,}$$

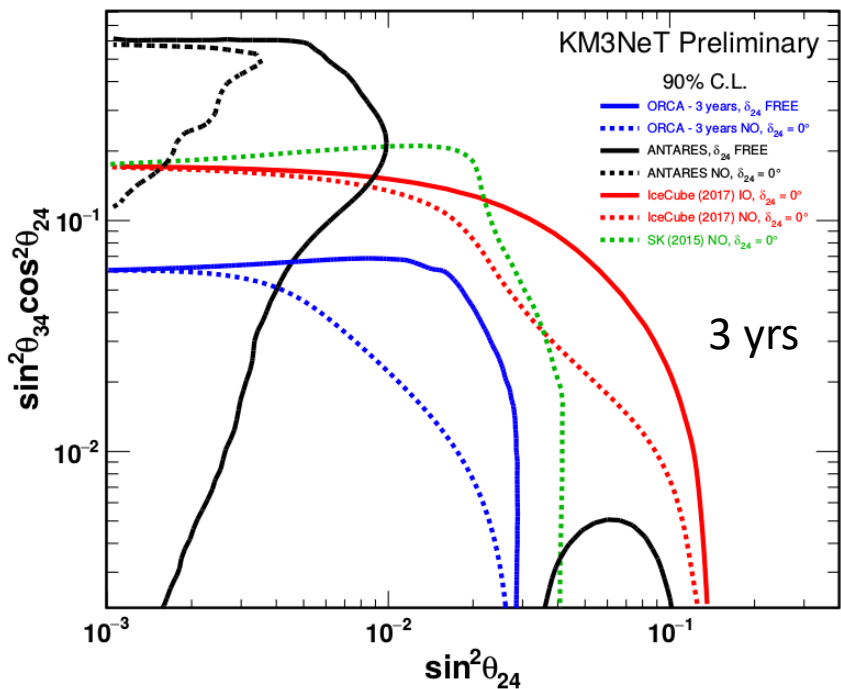
$$|U_{\tau 4}|^2 < 0.40 \text{ (0.68) at 90\% (99\%) CL.}$$



# KM3NeT/ORCA: sterile neutrino (3+1)

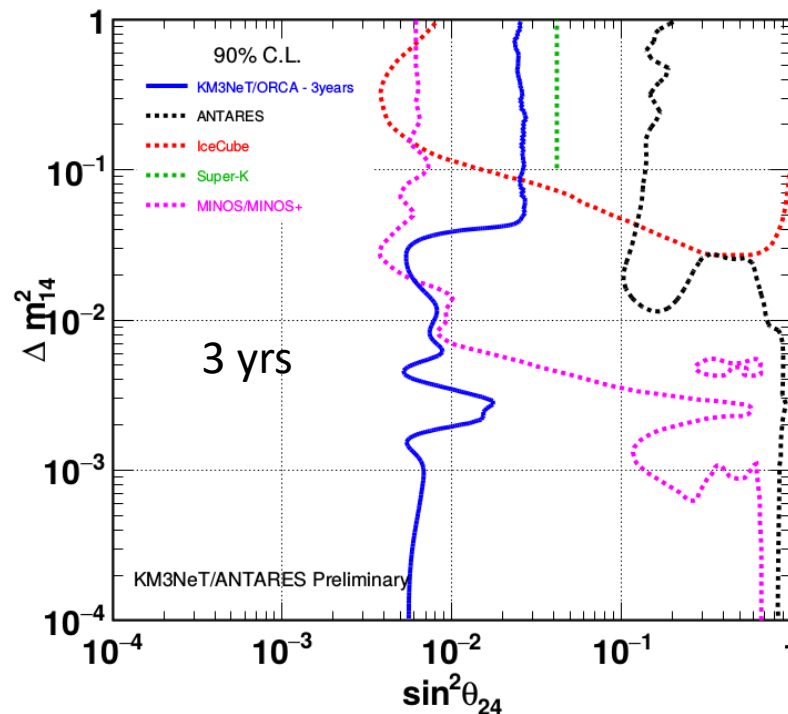
$$\Delta m_{41}^2 > 0.1 \text{ eV}^2$$

$$\Delta m_{41}^2 < 0.1 \text{ eV}^2$$



Dependence on  $\delta_{24}$

Factor of two better sensitivity on  $U_{\tau 4}$  than current limits from SK and IC

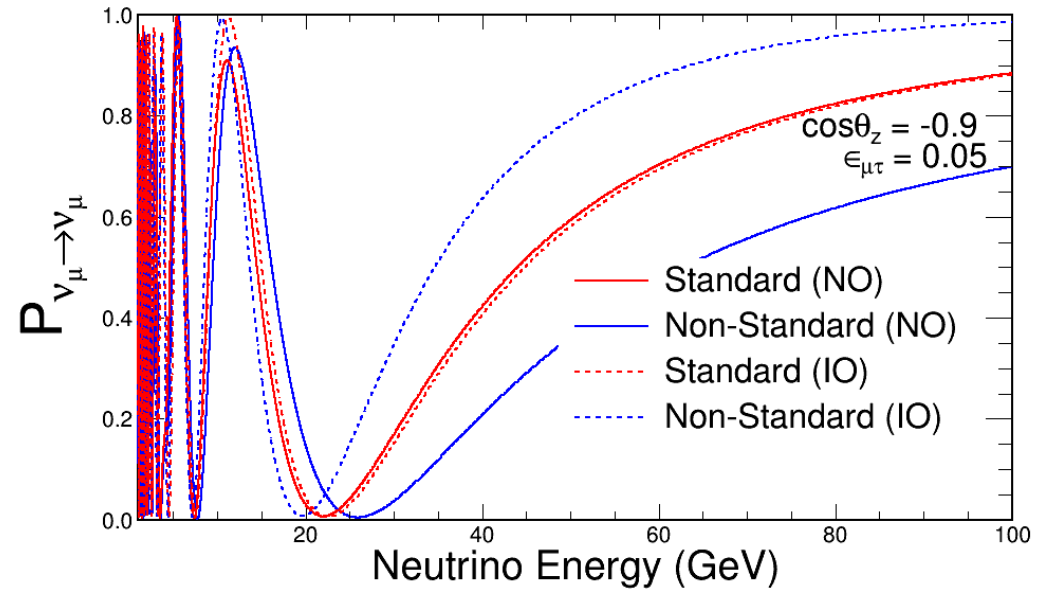


Due to longer & multiple baselines improve on MINOS/MINOS+ limits by 2 orders of magnitude



# KM3NeT/ORCA: non-standard interactions

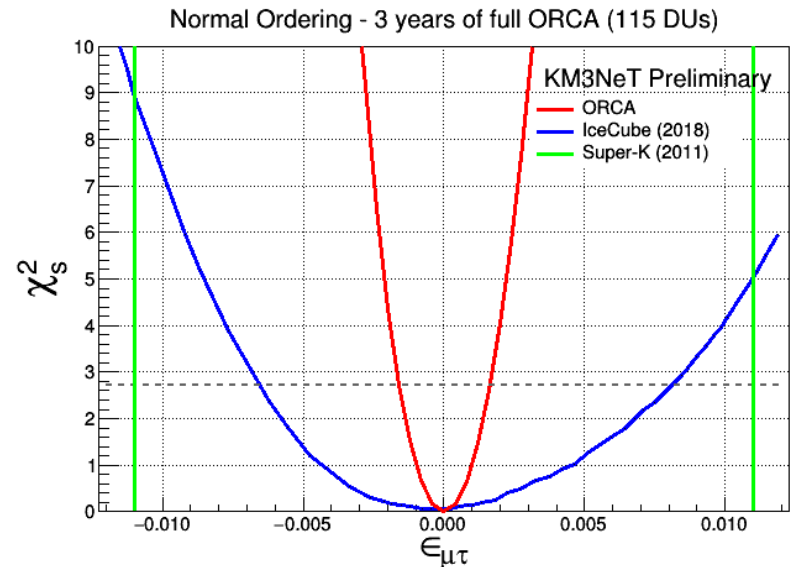
- ORCA sensitive to NSI effects of order 10% of the Fermi int.

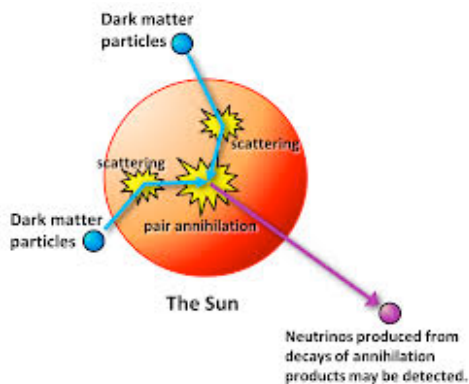


- Two-flavour hybrid model:

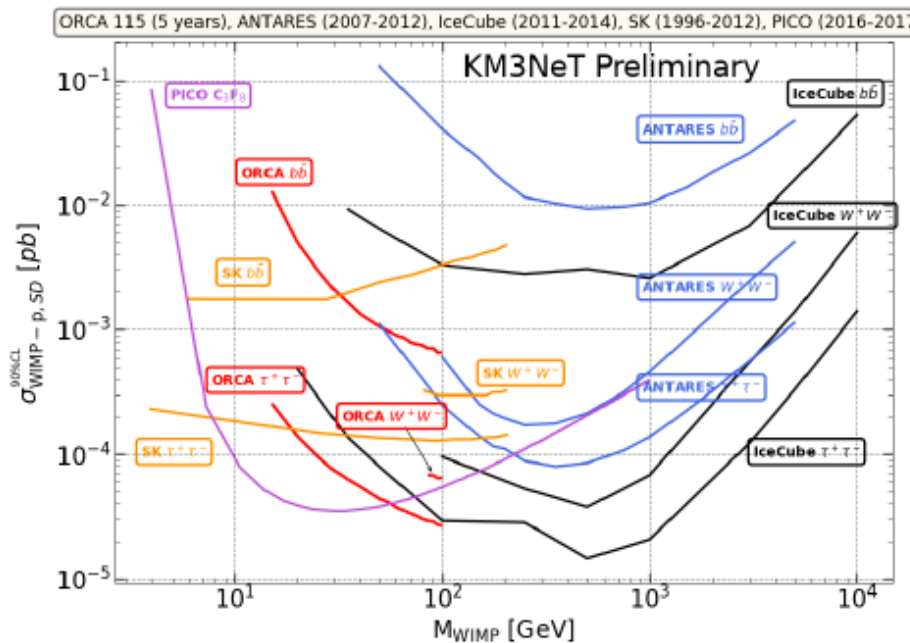
$$\epsilon_{\mu\mu} = \epsilon_{\tau\tau} = 0$$

- ORCA improves significantly over current atmospheric bounds

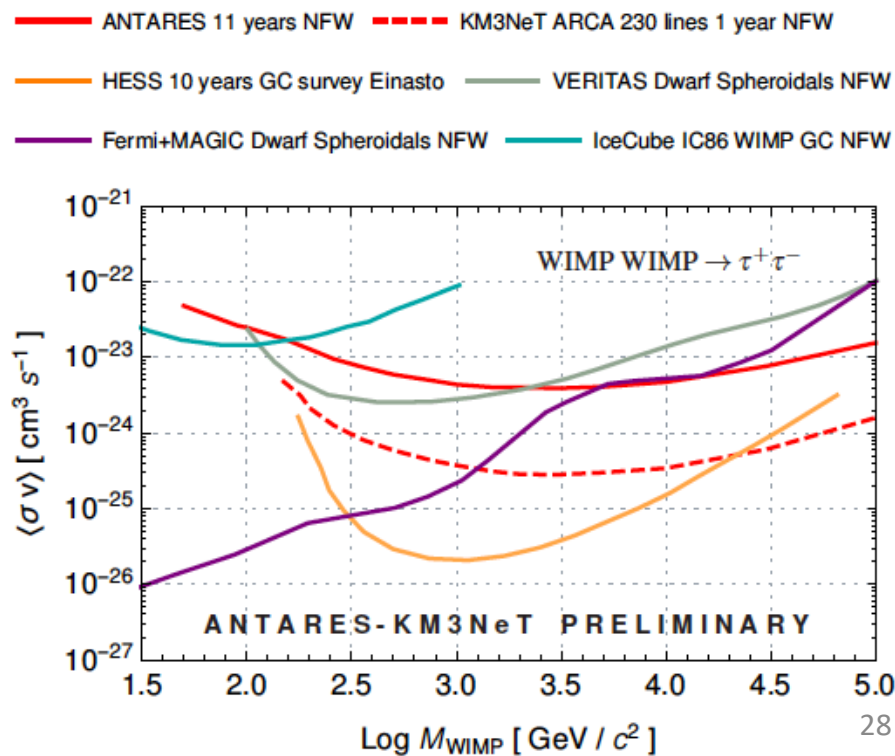


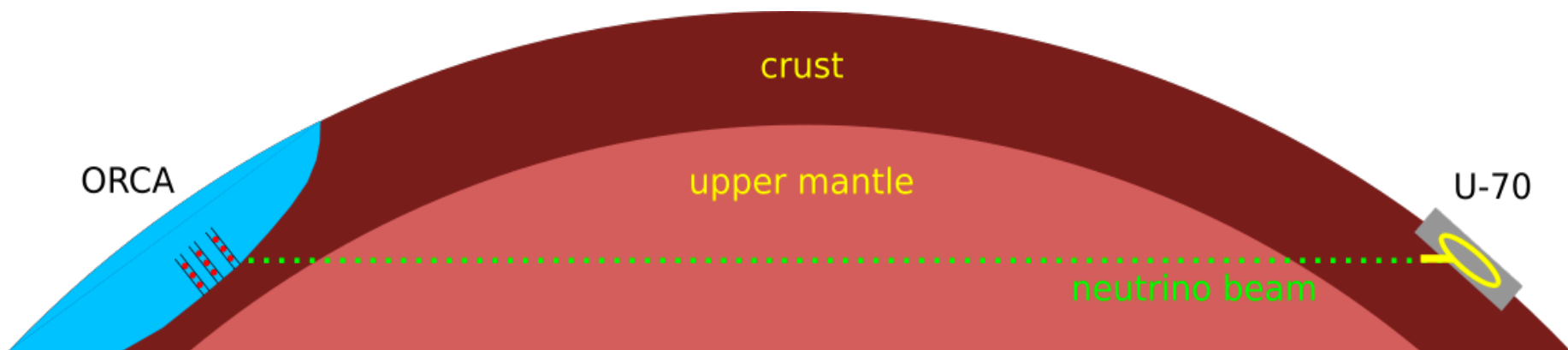


## Sun: ORCA115



## Galactic Centre: ARCA230

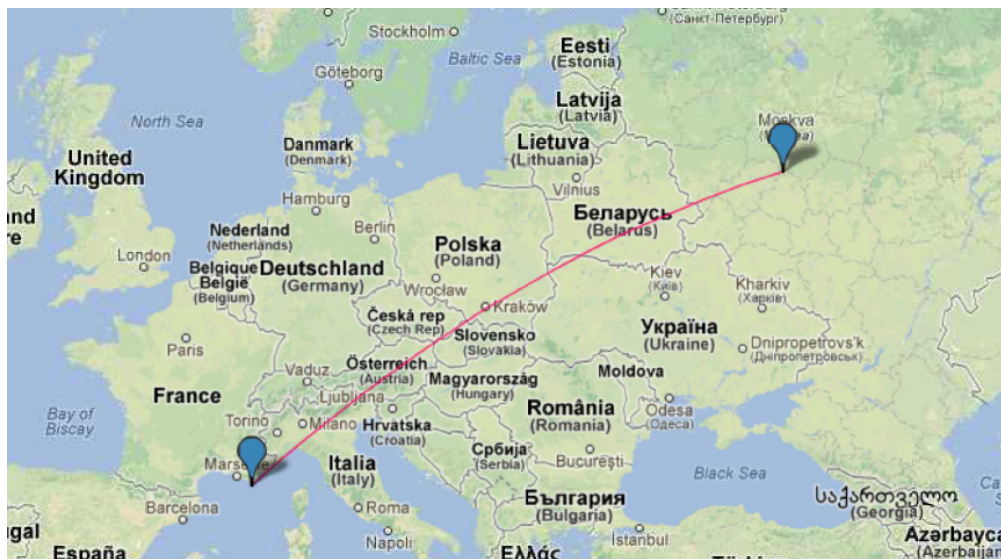




Big detector -> lower beam power

- Baseline 2588 km
- Beam inclination :  $11.7^\circ$  ( $\cos \theta = 0.2$ )
- Deepest point : 134 km ( $3.4 \text{ g/cm}^3$ )
- First oscillation maximum 5.1 GeV

-> Sensitivity to mass hierarchy and CP violation

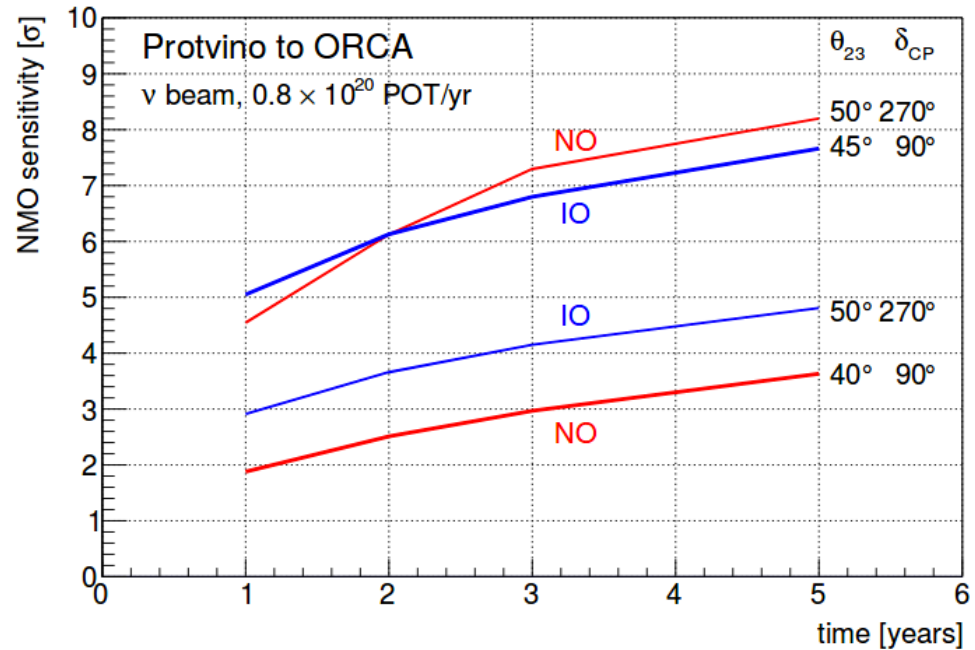




# Protvino to ORCA (P2O): prelim. study

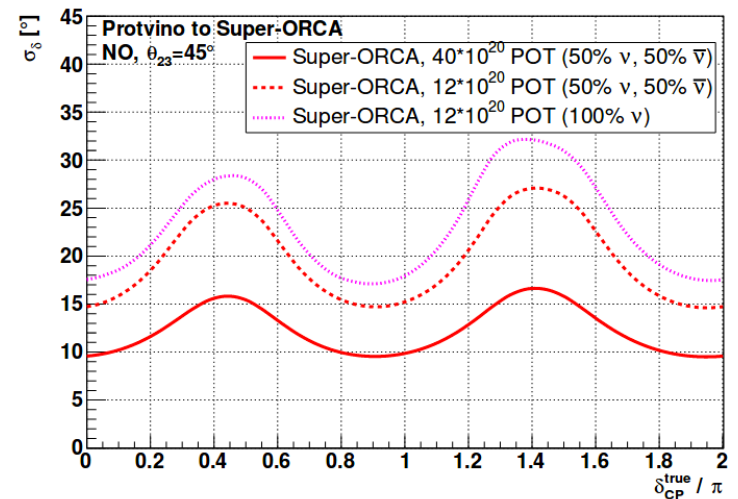
## Phase 1: Mass Hierarchy

ORCA detector +  
1 (5) years with 450 (90) kW



## Phase 2: CP Violation

10x denser detector (Super ORCA)  
450 kW  
Measure CP Phase to 10-16 degrees in 10 years



# Summary and Perspectives

ANTARES: Demonstration of potential of deep sea neutrino telescopes

KM3NeT: phased construction of a next-generation neutrino telescope  
Developed novel and performant multi-PMT technology  
interest from IC-Gen2, CHIPs, NuPrism, HyperK,...

ARCA-high energy:

- unprecedented angular resolution/multi-flavour astronomy
- investigation of diffuse cosmic flux, galactic sources,...

ORCA-low energy:

- NMH at 3 sigma level in 3 years (IH, NH/first octant).  
Much quicker if NH/second octant
- Competitive measurements of  $\Delta m^2_{32}$  and  $\sin^2\theta_{23}$ , tau appearance, sterile neutrinos, NSI, DM, tomography,...

CP Violation?:

- P2O: Protvino beam to (Super) ORCA

Exciting times ahead- please come and join us!





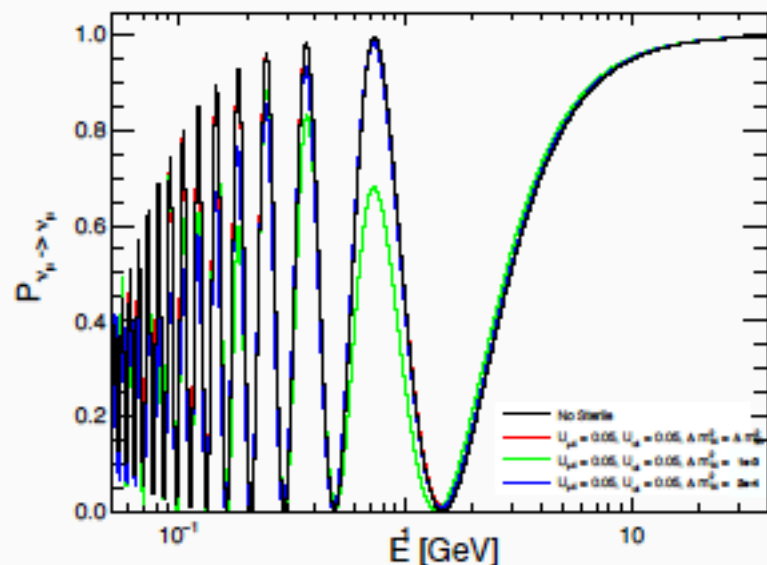
# Thanks!



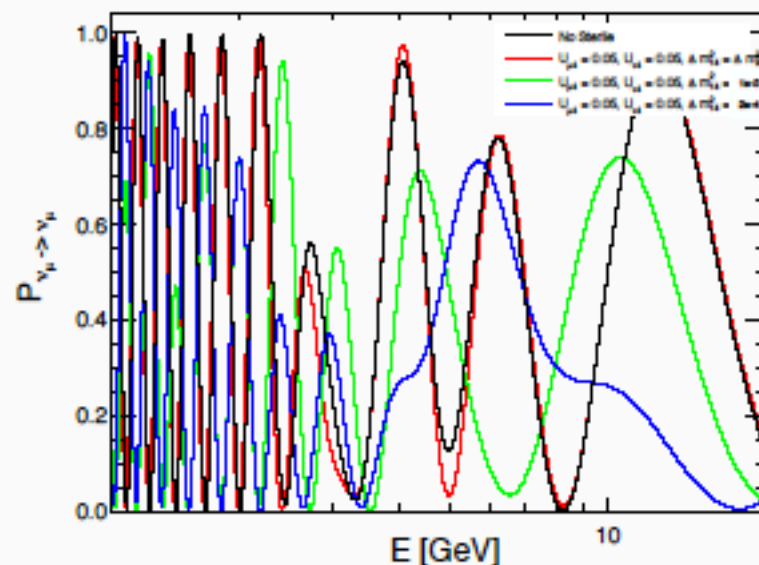
# BACKUPS

# ORCA vs MINOS

MINOS baseline + Vacuum



ORCA baseline + Matter



- We should expect a better sensitivity of ORCA wrt MINOS in the low sterile mass range.



# ORCA Seafloor

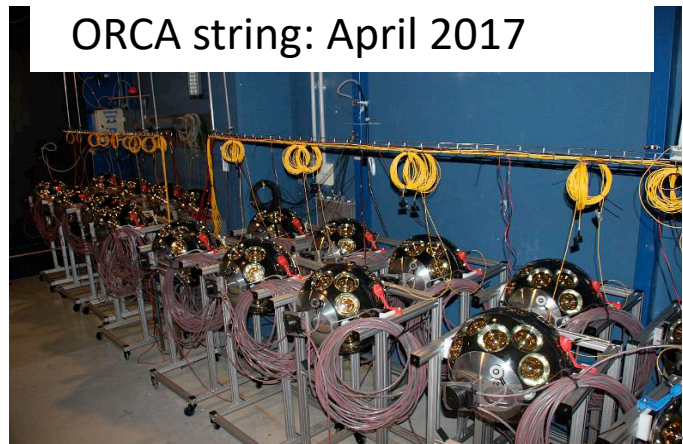
MEOC: Dec 2015, March 2017



1<sup>st</sup> node: Sept 2016, Sept 2018

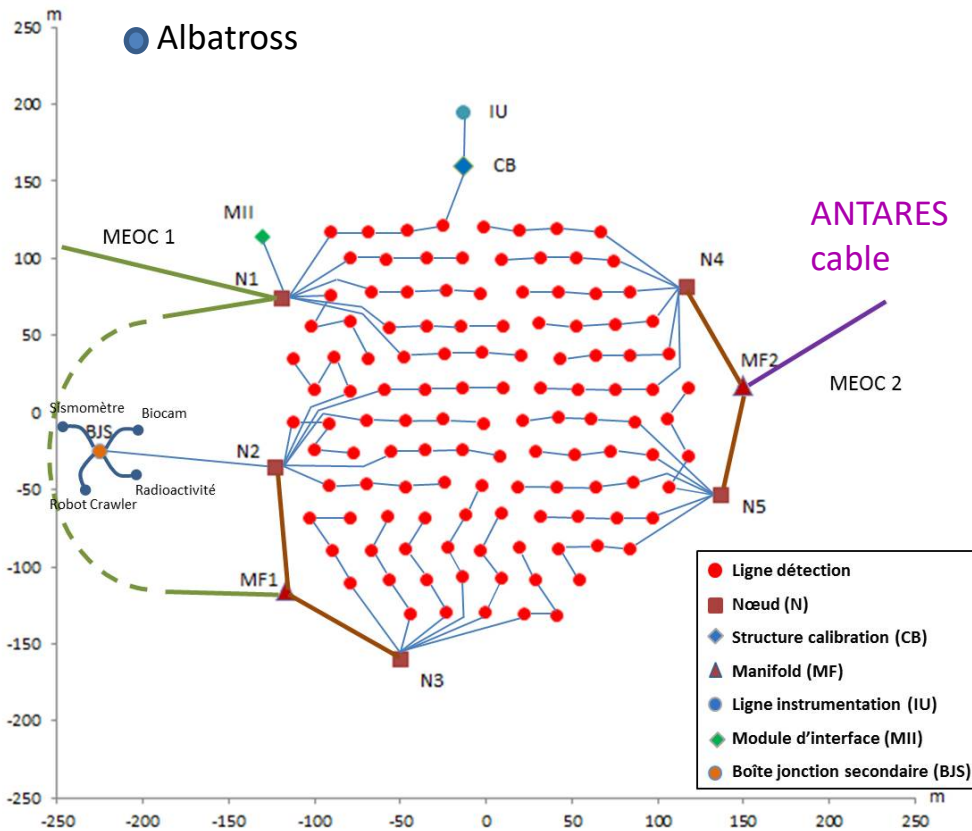


ORCA string: April 2017

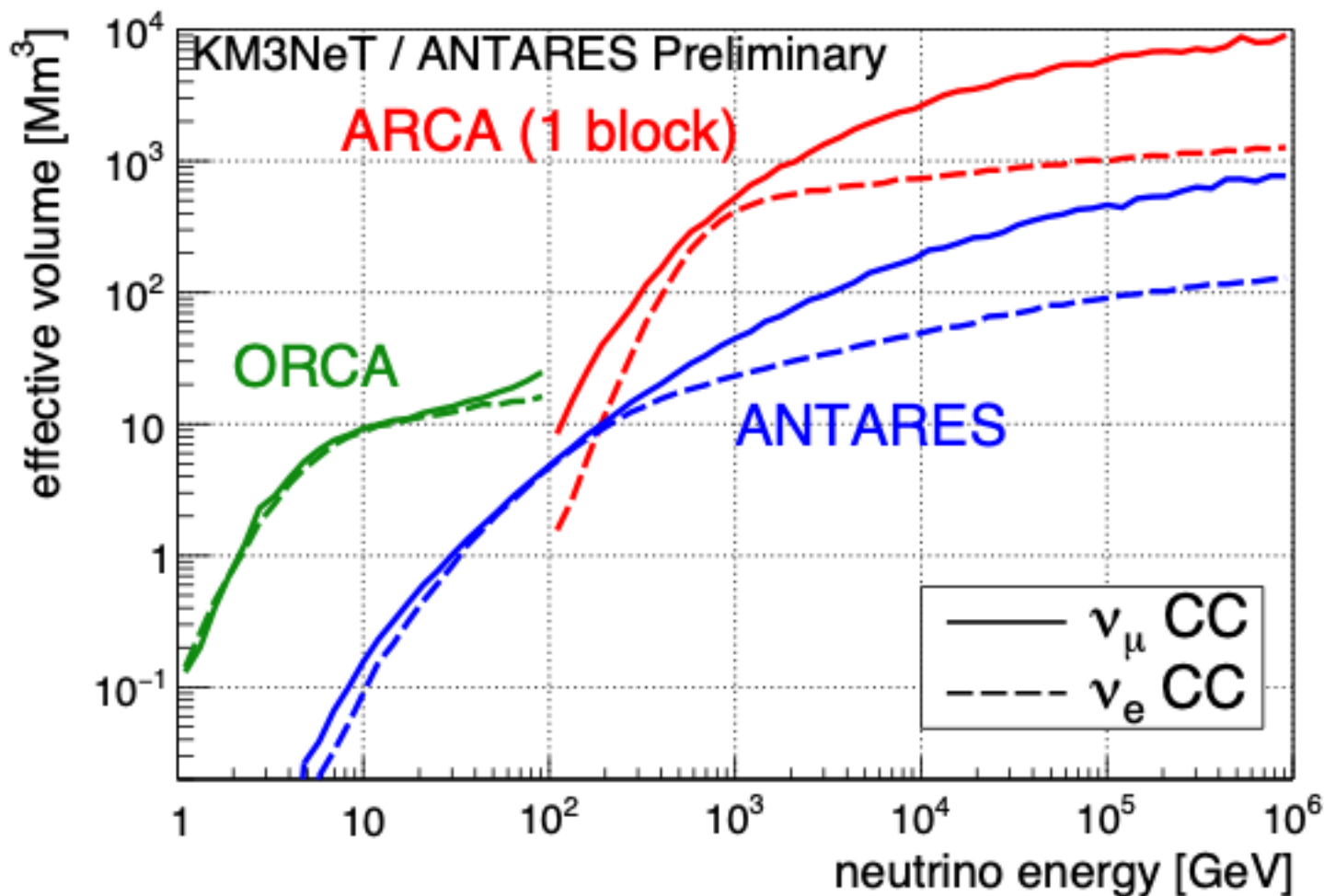


**Phase 1:** 6 string array at KM3NeT-France site to demonstrate technology/detection methods in the GeV range

**Phase 2:** Deploy 1 building block (115 strings)-2024



# ANTARES/ORCA/ARCA: Effective volumes



approved: ICRC2017, S. Hallmann

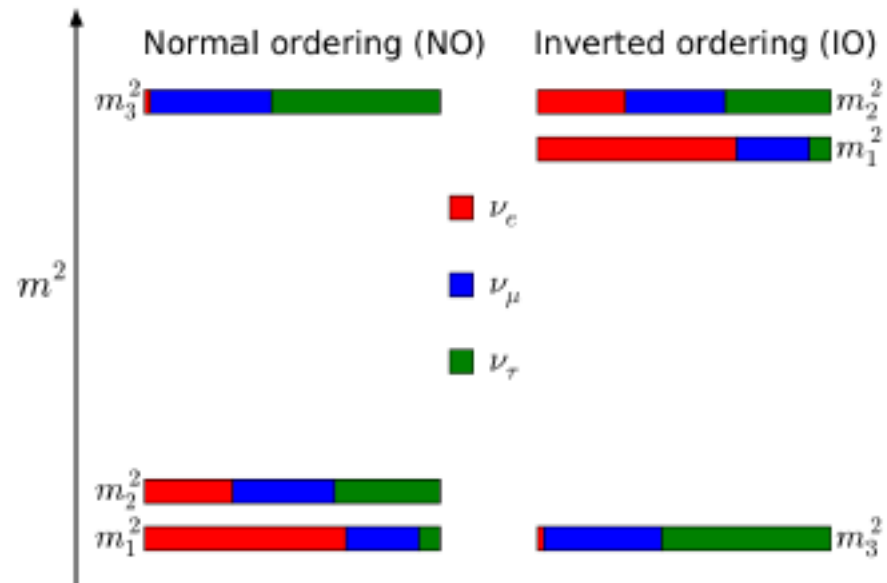
# Oscillation of massive neutrinos

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmospheric } \theta_A \sim 45^\circ} \cdot \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix}}_{\text{Reactor } \theta_{13} \sim 9^\circ} \cdot \underbrace{\begin{pmatrix} c_{21} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar } \theta_\odot \sim 30^\circ} \cdot \underbrace{\begin{pmatrix} e^{i\eta_1} & 0 & 0 \\ 0 & e^{i\eta_2} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Majorana}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$\downarrow$   
 CP violating phase  $\delta_{CP}$

All parameters measured to fair precision except:

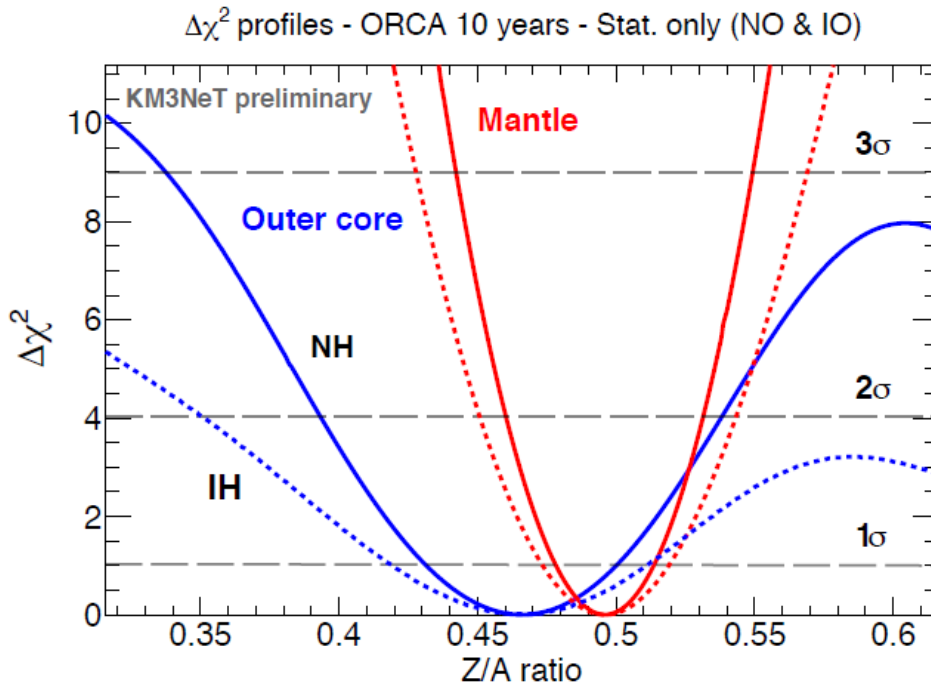
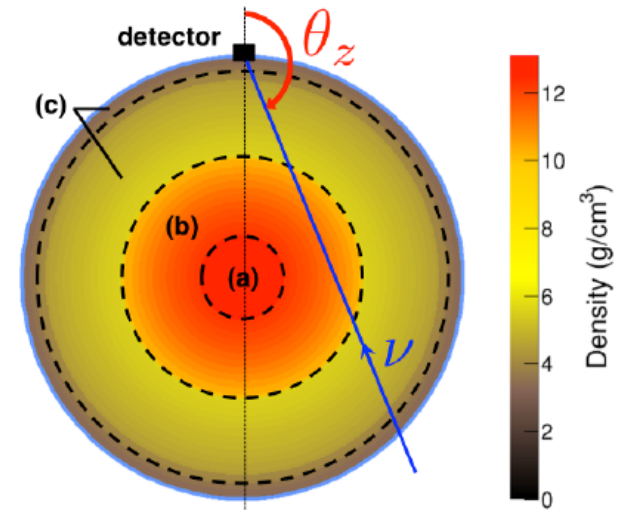
mass ordering  
 octant of  $\theta_{23}$   
 CP phase  
 Absolute masses  
 Majorana/Dirac



$$m_1^2 < m_2^2 \\
 m_2^2 - m_1^2 \ll |m_3^2 - m_{1,2}^2|$$

# Earth Tomography

- ORCA is sensitive to the electron density  $N_e$  while geophysics measure  $\rho_m$
- $1\sigma$  stat. uncertainty after 10 years for NH:
  - ~ 4% in the whole mantle (c)
  - ~ 7% in the whole outer core (b)



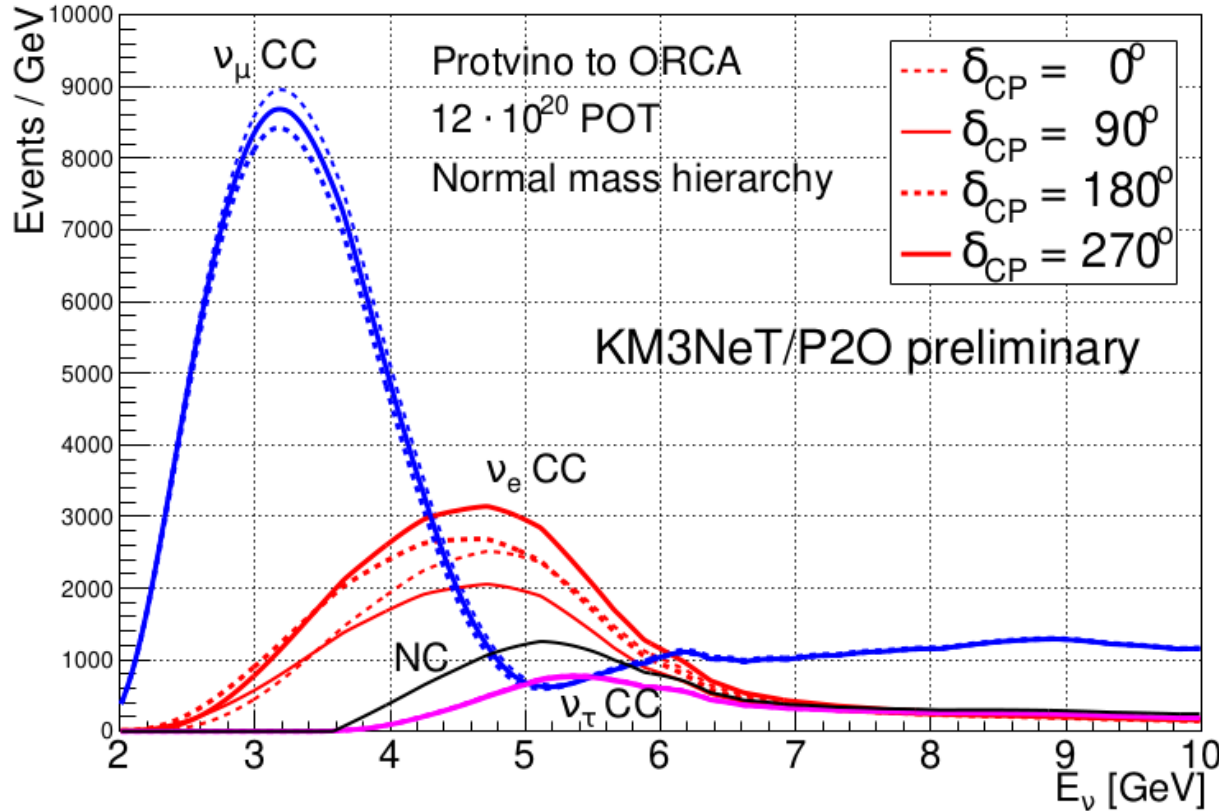
S Bourret et al (KM3NeT) 2016  
*J. Phys.: Conf. Ser. [Neutrino2016]*

$$\frac{N_e}{\rho_m} \propto \sum_i w_i \frac{Z_i}{A_i}$$

- PREM model basis for  $\rho_m$
- uniform  $Z/A$  rescaling in layer
- Monte Carlo response & PID
- statistical uncertainty only



# P2O: Expected rates in ORCA (NH)



After 3 yr of 450 kW beam:

- $\nu_\mu$  CC: ~ 30000 events
- $\nu_e$  CC: ~ 8000 events
- $\nu_\tau$  CC: ~ 3500 events
- NC: ~ 6000 events

For comparison:

DUNE: ~ 750  $\nu_e$  / 3 yr

Vacuum oscillation maximum at  $E = 5.1$  GeV

Most  $\nu_\mu$  convert to  $\nu_\tau$  which remains largely invisible (CC reaction suppressed by  $\tau$  mass)

$\nu_\mu \rightarrow \nu_e$  transitions are enhanced by the matter effect, resonance energy 3.8 GeV